

A Tale of Two Anomalies: from LHCb to ANITA

Speaker: Yicong Sui

In collaboration with:
Wolfgang Altmannshofer, Bhupal Dev, Amarjit Soni



Introduction of B-anomaly



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$$R_D = \frac{\mathcal{B}(B \rightarrow D\tau\bar{\nu})}{\mathcal{B}(B \rightarrow D\ell\bar{\nu})} \quad R_{D^*} = \frac{\mathcal{B}(B \rightarrow D^*\tau\bar{\nu})}{\mathcal{B}(B \rightarrow D^*\ell\bar{\nu})}$$

BaBar, Belle : $\ell = e, \mu$
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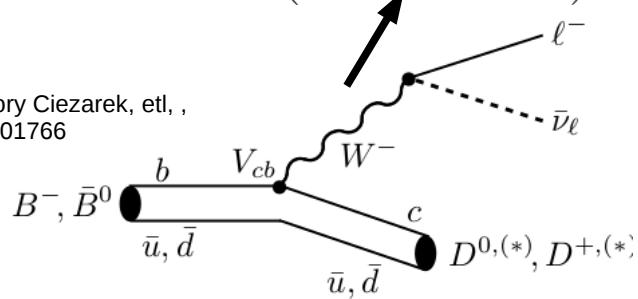
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Gregory Ciezarek, et al.,
1703.01766



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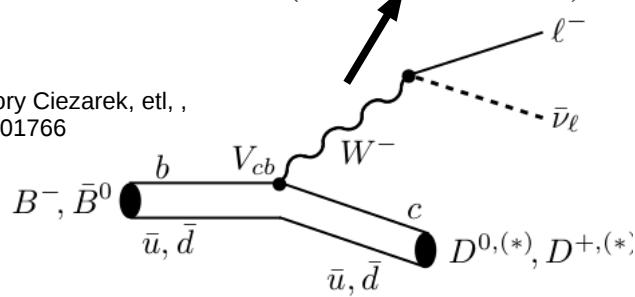
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Florian U. Bernlochner, et al, 1703.05330

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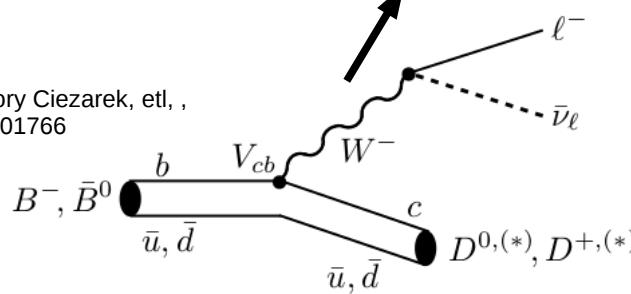
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PDG2018

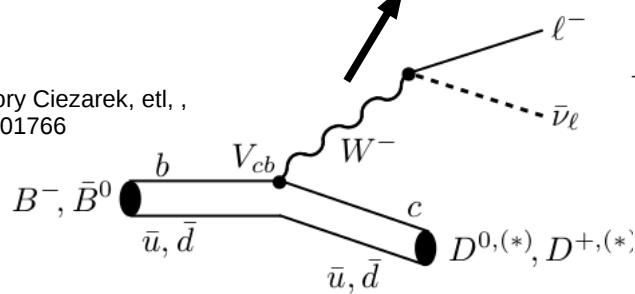
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PDG2018

2.4σ

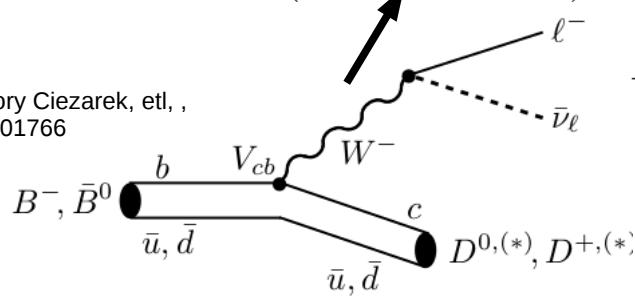
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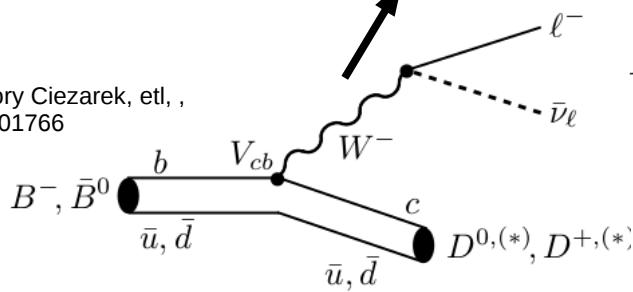
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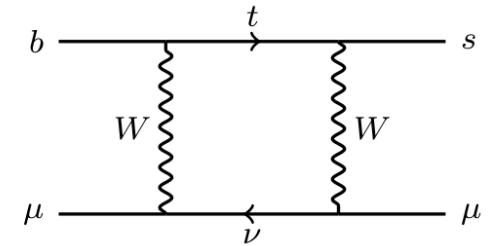
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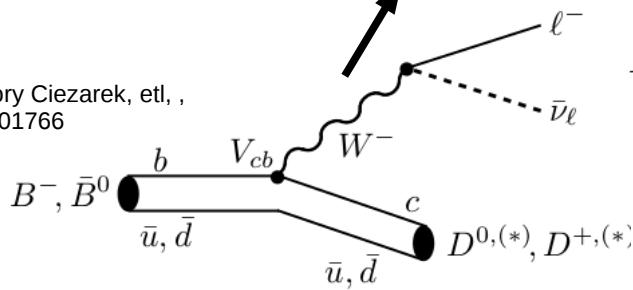
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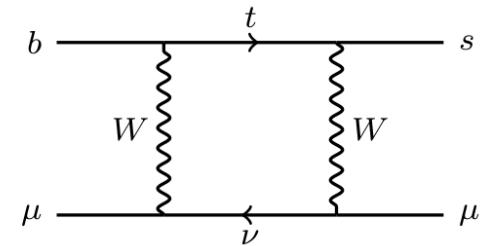
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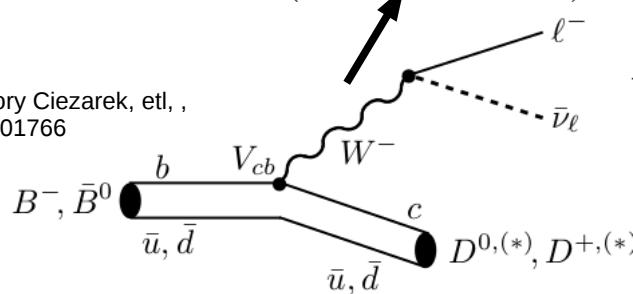
$$R_{K^*}^{\text{SM}}|_{q^2 \in [0.045, 1.1] \text{ GeV}} = 0.906 \pm 0.028. \quad R_{K^*}^{\text{SM}}|_{q^2 \in [1.1, 6] \text{ GeV}} = R_K^{\text{SM}} = 1.00 \pm 0.01.$$

Marzia Bordone, Gino Isidori, Andrea Pattori, 1605.07633

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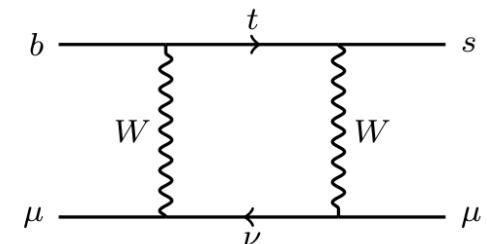
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PDG2018

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2.4σ



$$R_K = 0.846^{+0.060}_{-0.054}{}^{+0.016}_{-0.014},$$

for $1.1 \text{ GeV}^2 < q^2 < 6 \text{ GeV}^2$

LHCb collaboration, 1903.09252

$$R_{K^*} = \begin{cases} 0.66^{+0.11}_{-0.07} \pm 0.03, & \text{for } 0.045 \text{ GeV}^2 < q^2 < 1.1 \text{ GeV}^2, \\ 0.69^{+0.11}_{-0.07} \pm 0.05, & \text{for } 1.1 \text{ GeV}^2 < q^2 < 6 \text{ GeV}^2, \end{cases}$$

LHCb collaboration, 1705.05802

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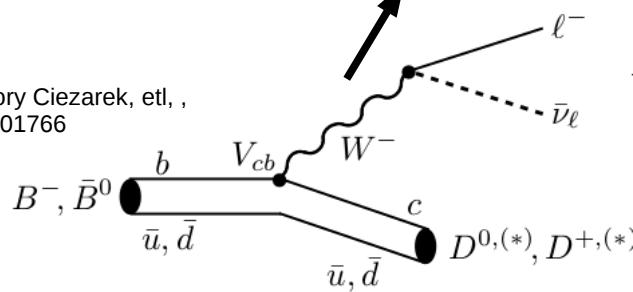
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$$2.5\sigma$$

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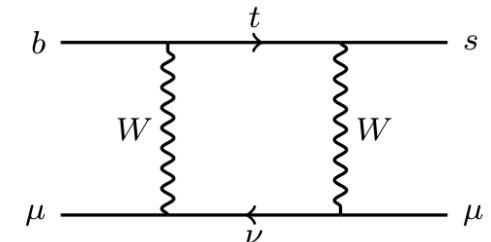
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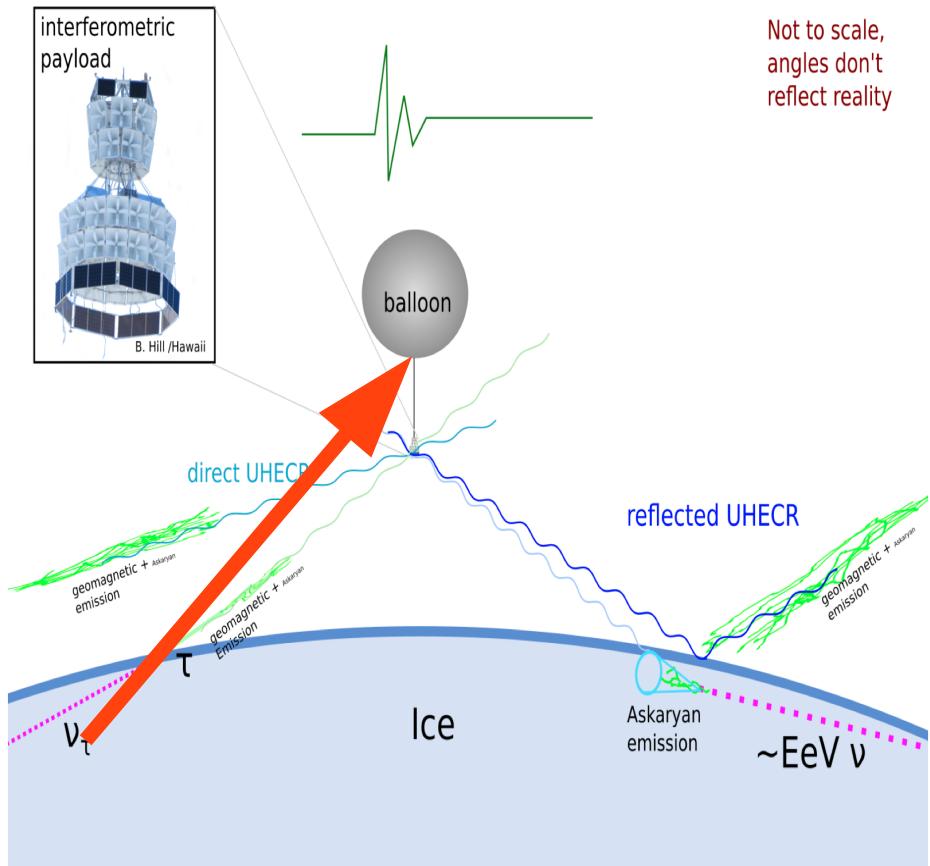
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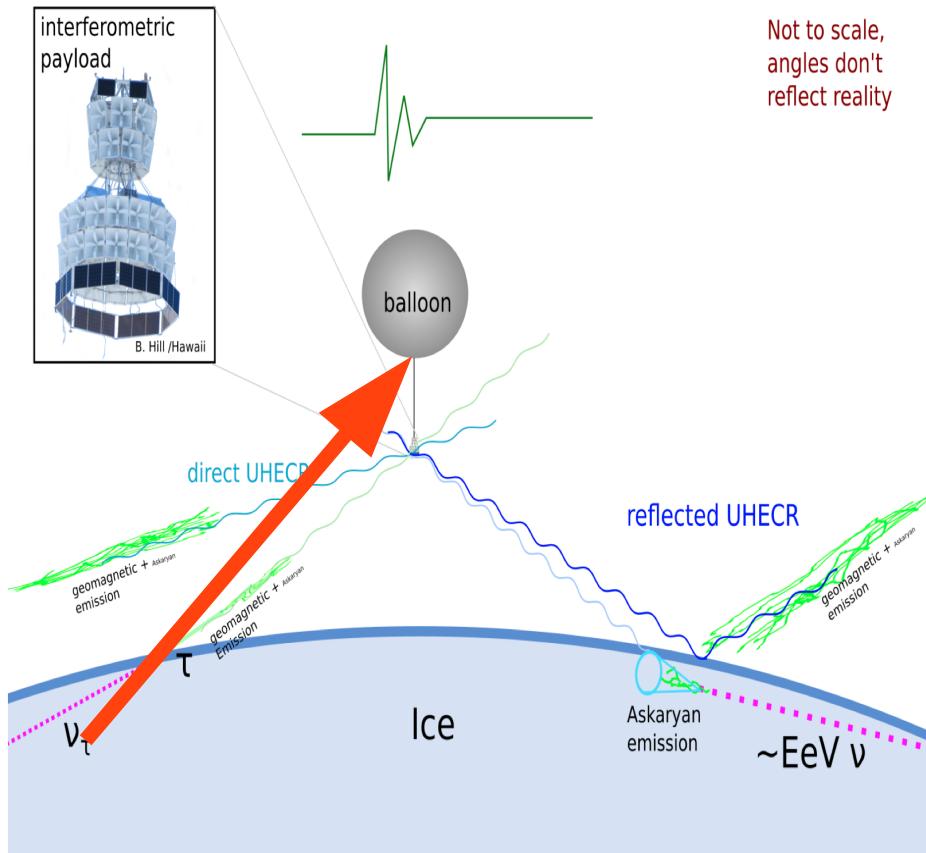
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The ANITA detection concepts, figure from Cosmin Deaconu



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TABLE I: ANITA-I,-III anomalous upward air showers.

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date, time	2006-12-28,00:33:20UTC	2014-12-20,08:33:22.5UTC
Lat., Lon. ⁽¹⁾	-82.6559, 17.2842	-81.39856, 129.01626
Altitude	2.56 km	2.75 km
Ice depth	3.53 km	3.22 km
El., Az.	$-27.4 \pm 0.3^\circ, 159.62 \pm 0.7^\circ$	$-35.0 \pm 0.3^\circ, 61.41 \pm 0.7^\circ$
RA, Dec ⁽²⁾	$282.14064, +20.33043$	$50.78203, +38.65498$
$E_{\text{shower}}^{(3)}$	$0.6 \pm 0.4 \text{ EeV}$	$0.56_{-0.2}^{+0.3} \text{ EeV}$

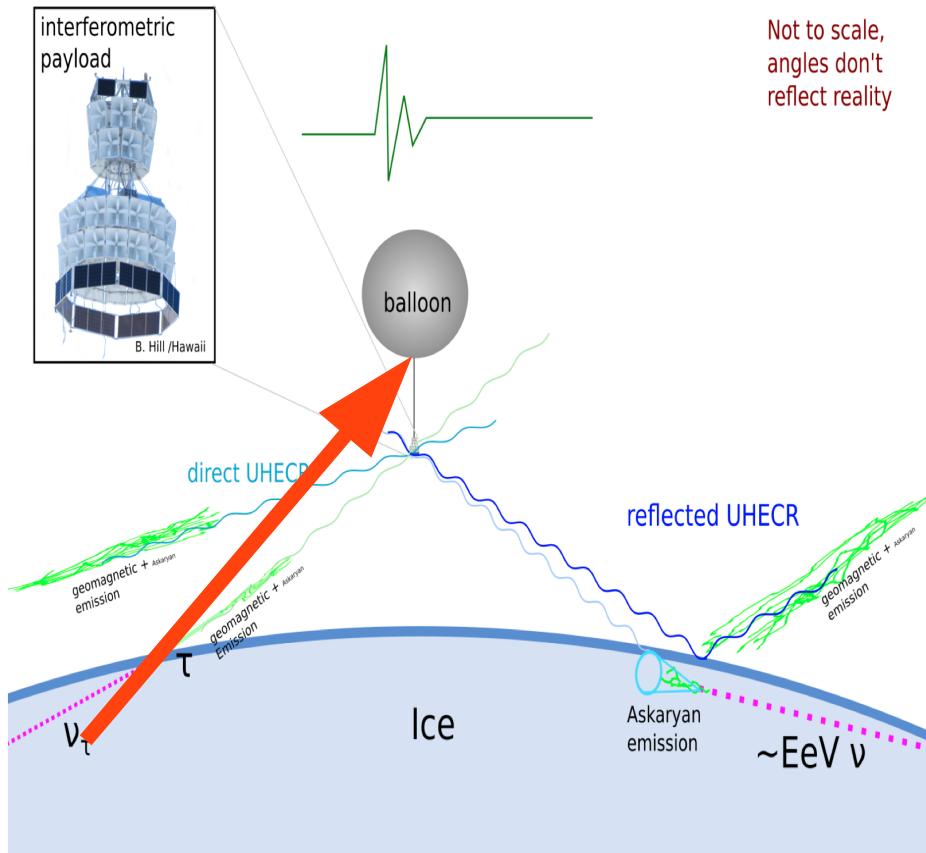
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² Sky coordinates projected from event arrival angles at ANITA.

³ For upward shower initiation at or near ice surface.

Table from ANITA, 1803.05088

Introduction of ANITA anomaly



The ANITA detection concepts, figure from Cosmin Deaconu

Properties of the anomalous upward events

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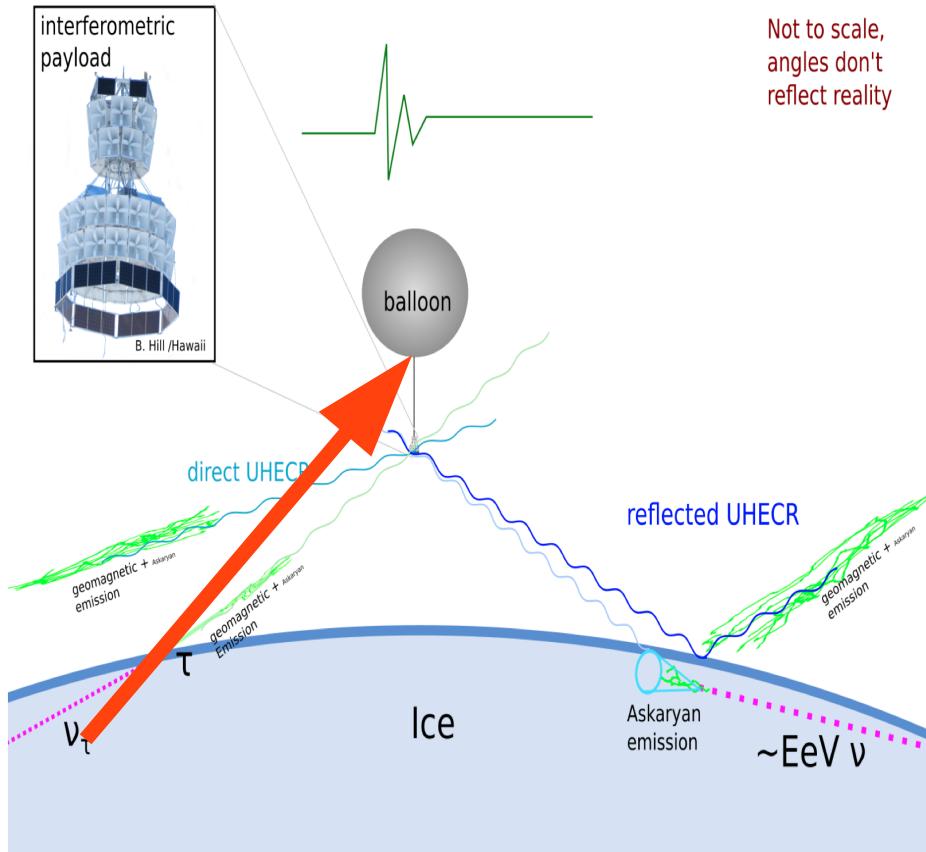
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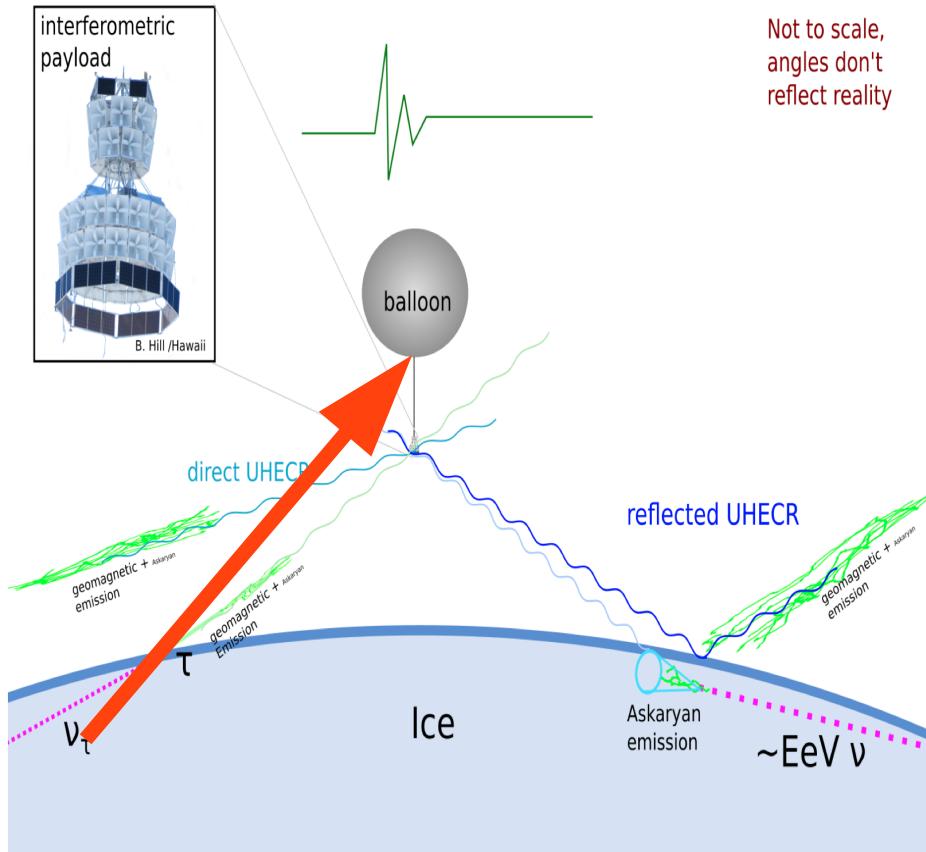
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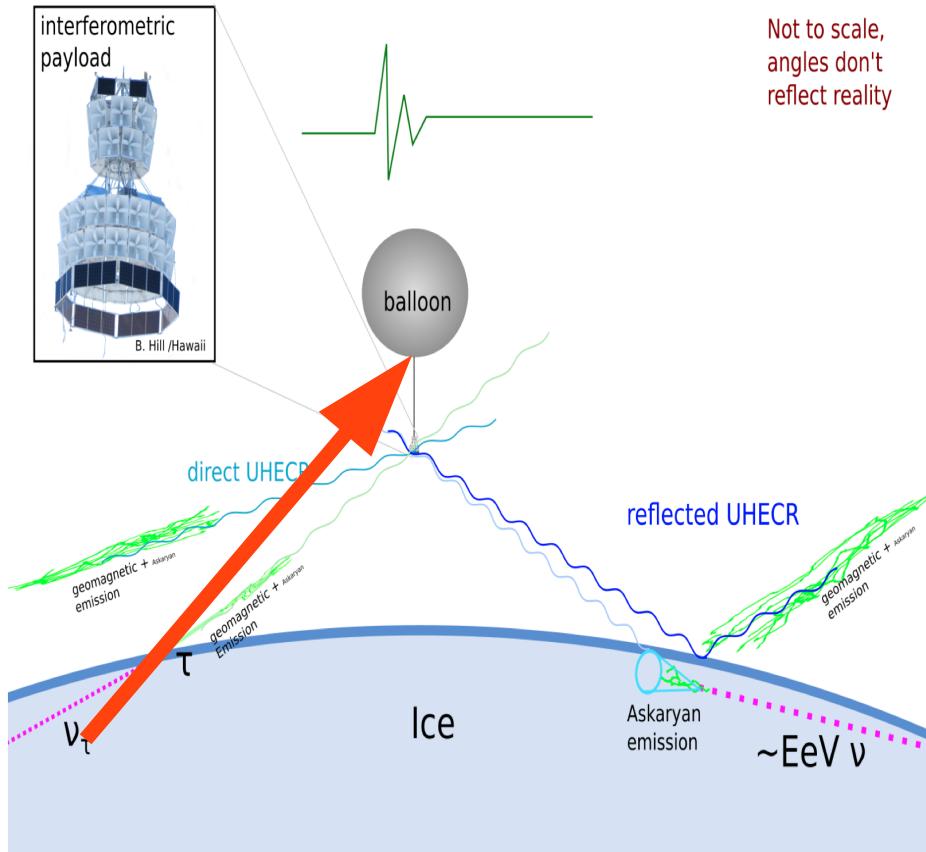
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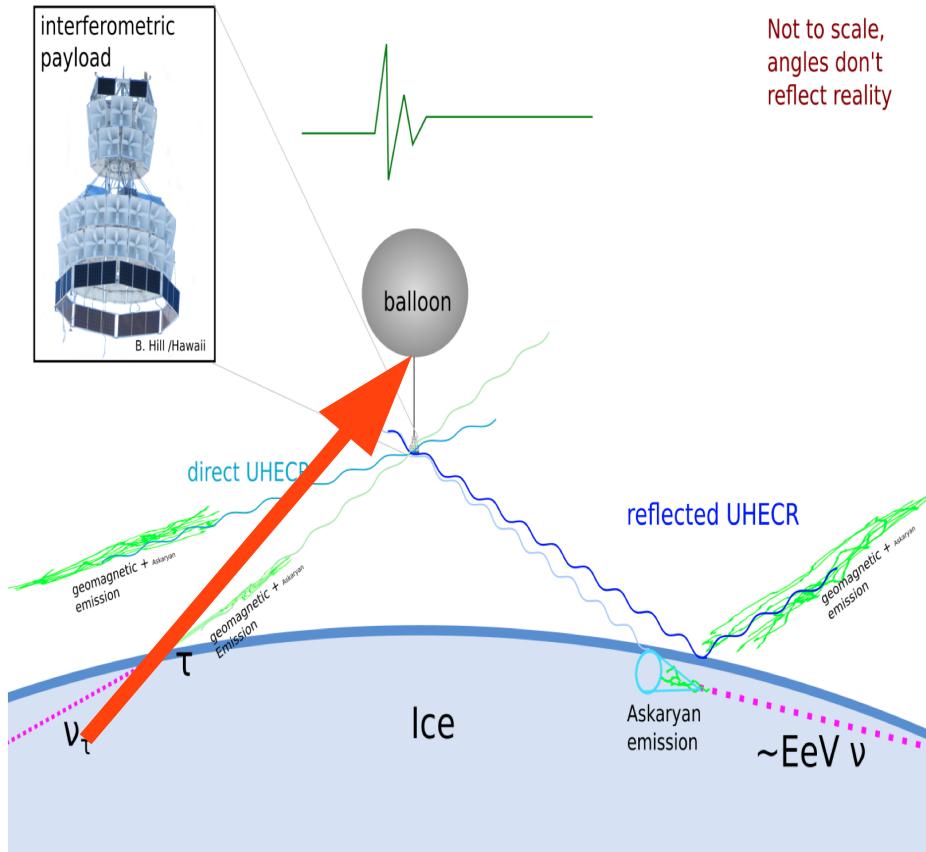
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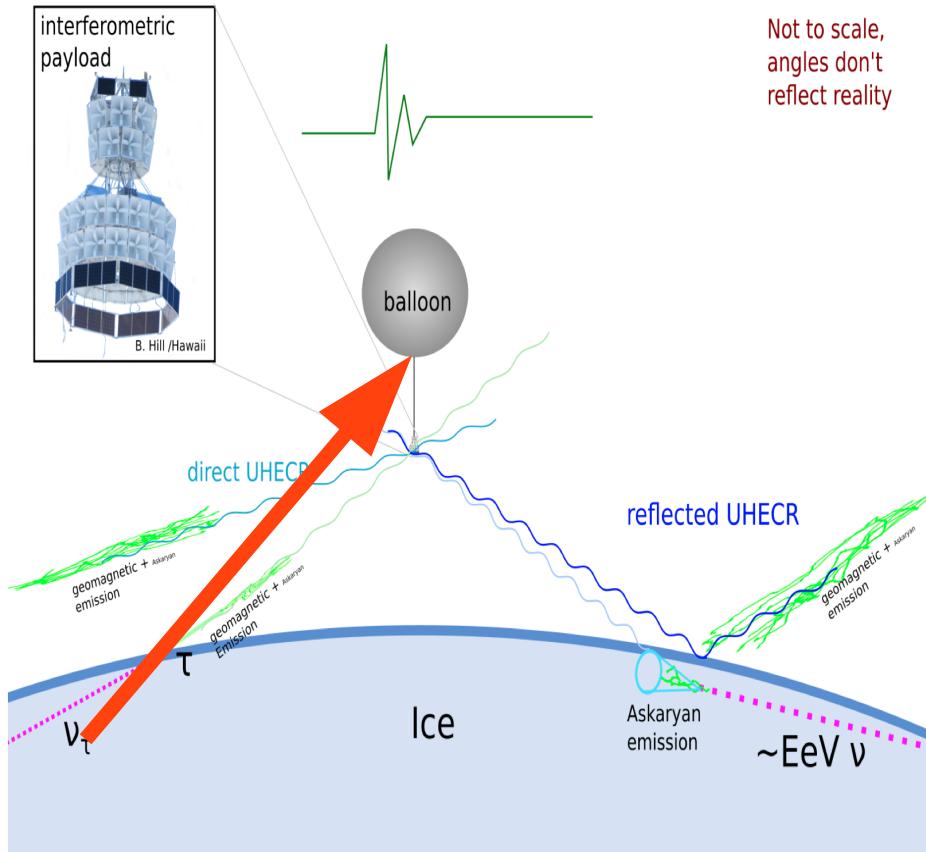
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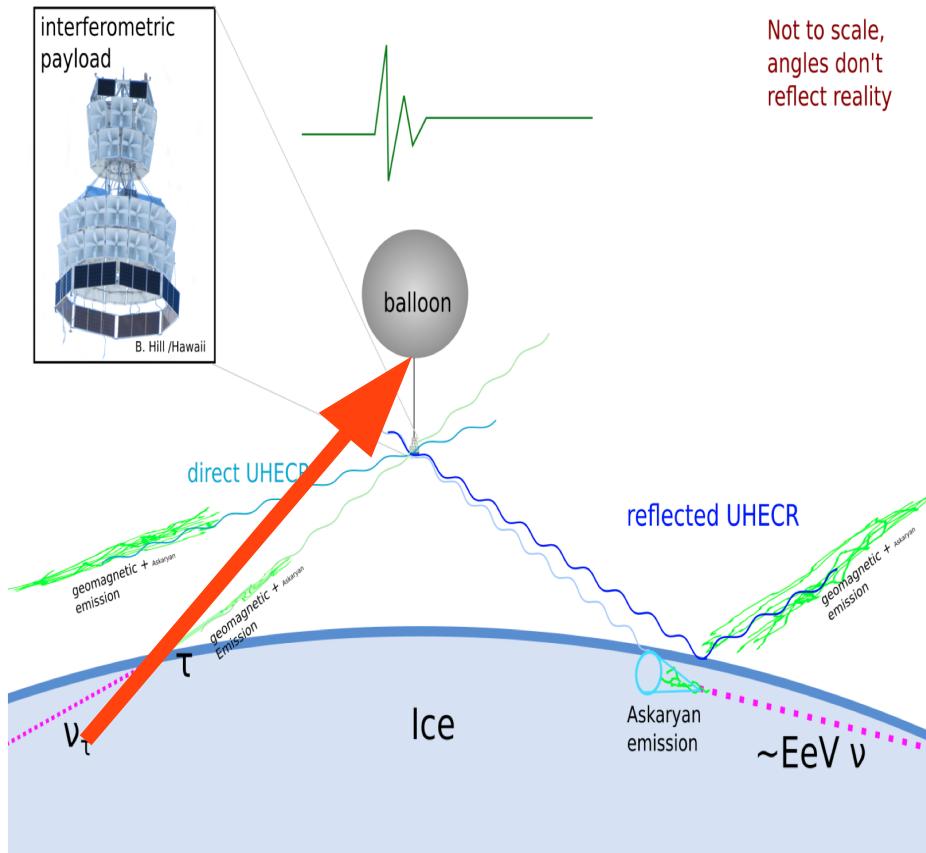
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$$\begin{aligned} R &\sim 6400 \text{ km} \\ D &= 2R \cos(\theta) > 5700 \text{ km} \\ l_{SM} &\sim 300 \text{ km (in rock)} \end{aligned}$$

TABLE I: ANITA-I,-III anomalous upward air showers.

event, flight	3985267, ANITA-I	15717147, ANITA-III
date, time	2006-12-28, 00:33:20 UTC	2014-12-20, 08:33:22.5 UTC
Lat., Lon. ⁽¹⁾	-82.6559, 17.2842	-81.39856, 129.01626
Altitude	2.56 km	2.75 km
Ice depth	3.53 km	3.22 km
El., Az.	$-27.4 \pm 0.3^\circ, 159.62 \pm 0.7^\circ$	$-35.0 \pm 0.3^\circ, 61.41 \pm 0.7^\circ$
RA, Dec ⁽²⁾	$282.14064, +20.33043$	$50.78203, +38.65498$
$E_{\text{shower}}^{(3)}$	$0.6 \pm 0.4 \text{ EeV}$	$0.56^{+0.3}_{-0.2} \text{ EeV}$

¹ Latitude, Longitude of the estimated ground position of the event.

² Sky coordinates projected from event arrival angles at ANITA.

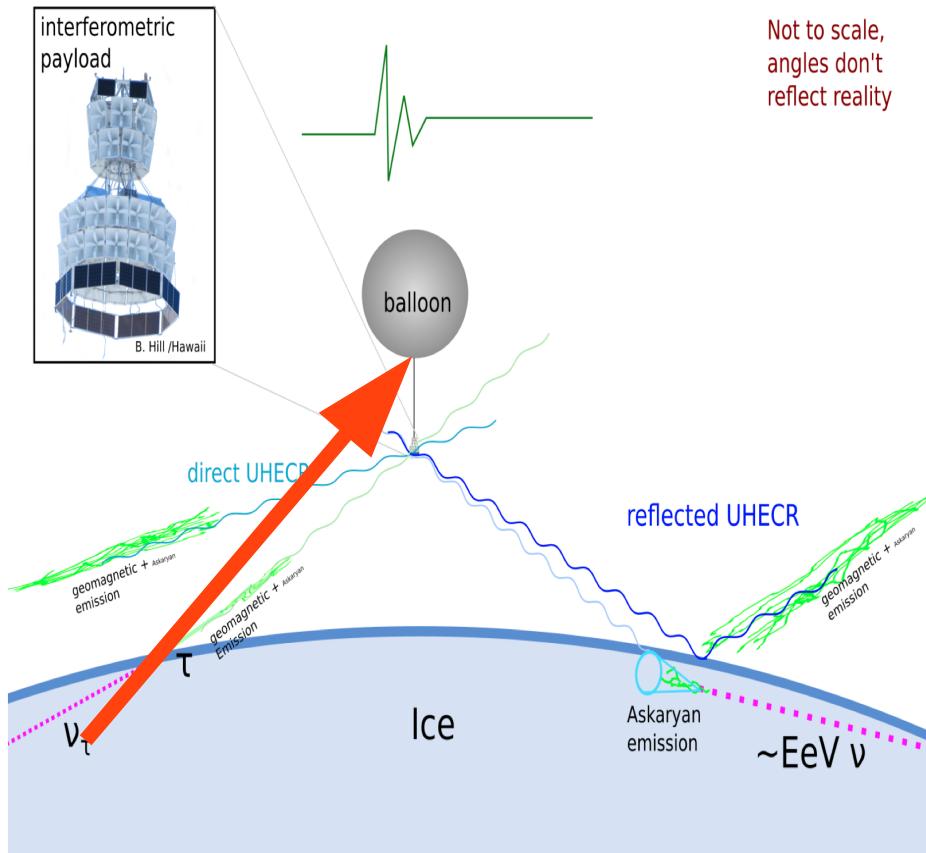
³ For upward shower initiation at or near ice surface.

Table from ANITA, 1803.05088



Introduction of ANITA anomaly

$$P_{\text{survival}} \sim 10^{-6}$$



The ANITA detection concepts, figure from Cosmin Deaconu

Properties of the anomalous upward events

1. Large Elevation Angle, going upwards.
2. No Polarity Reverse Relative to Geomagnetic Field.
3. Both have large energy~0.5EeV
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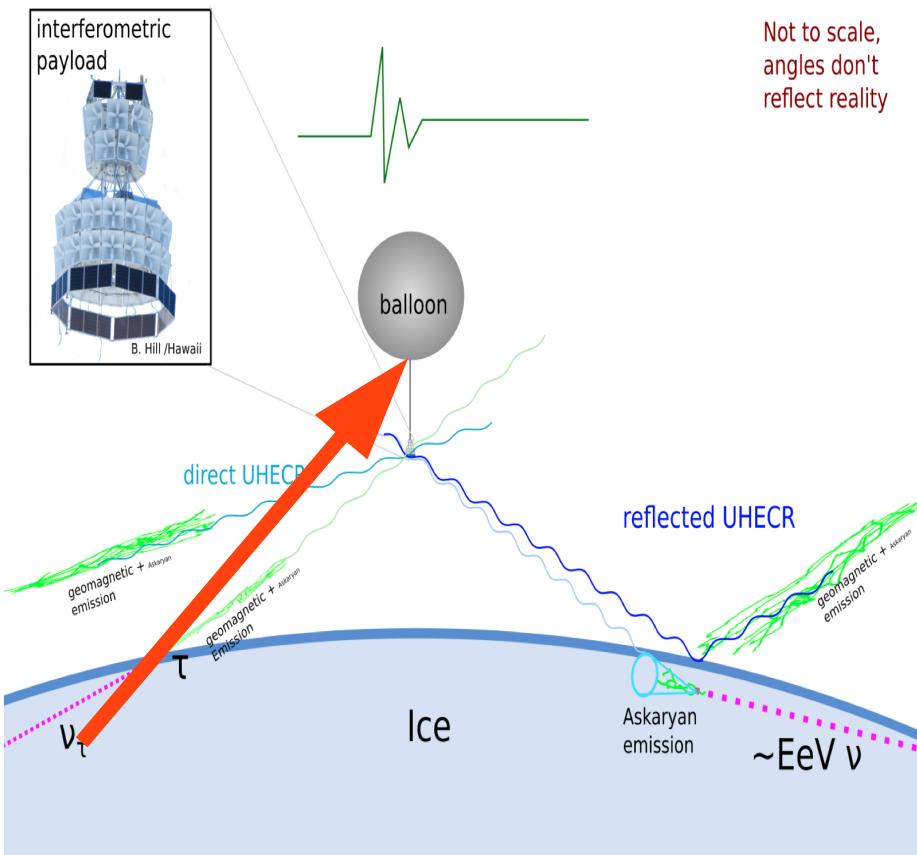
Table from ANITA, 1803.05088



Introduction of ANITA anomaly

$$P_{\text{survival}} \sim 10^{-6}$$

$$= 0.0001\%$$



The ANITA detection concepts, figure from Cosmin Deaconu

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Table from ANITA, 1803.05088



New Physics Interpretation to ANITA-anomaly

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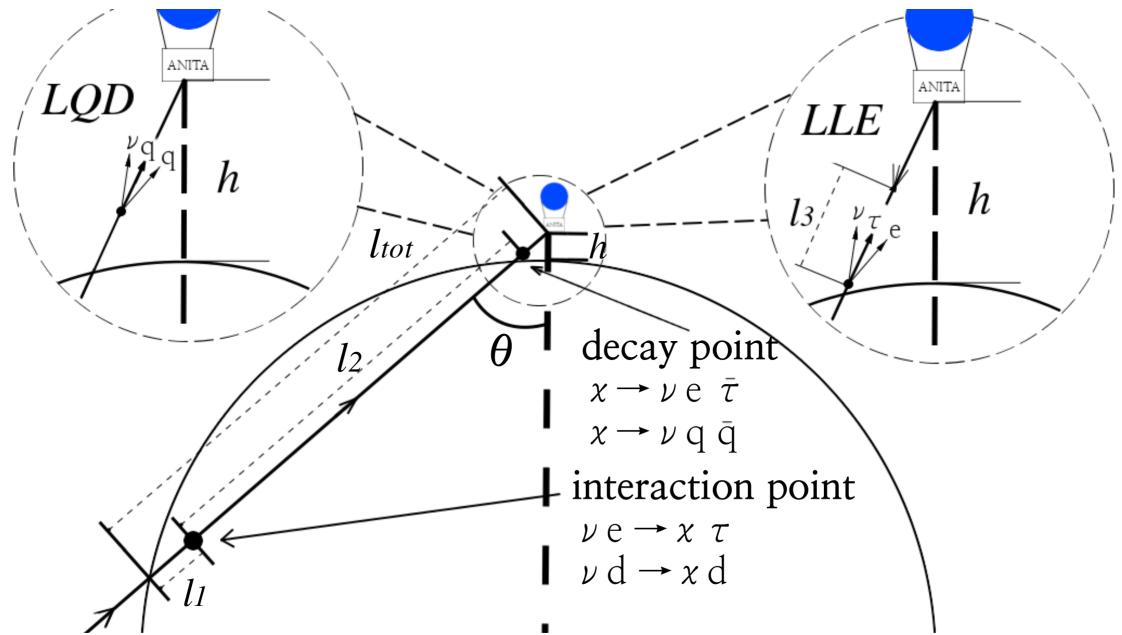
Long-lived neutral particle χ + TeV level mediator particle

Jack Collins and Bhupal Dev, **Yicong Sui**, 1810.08479

New Physics Interpretation to ANITA-anomaly

Long-lived neutral particle χ + TeV level mediator particle

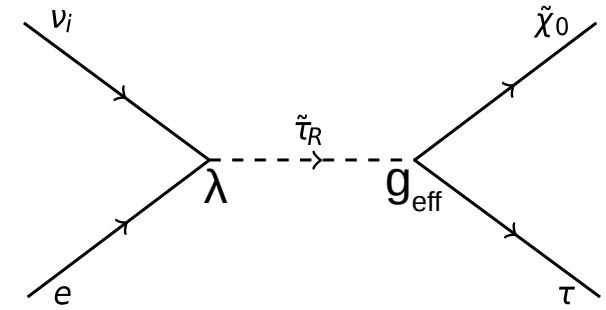
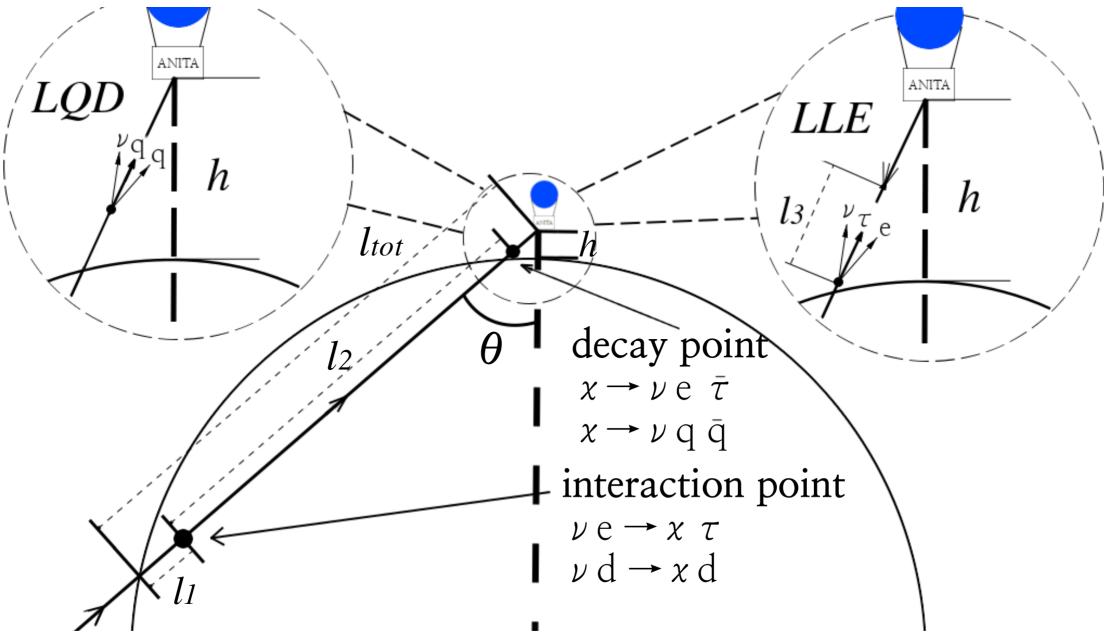
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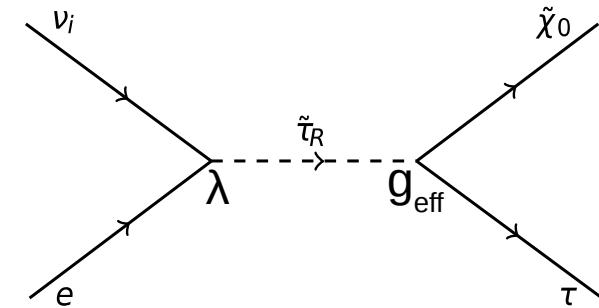
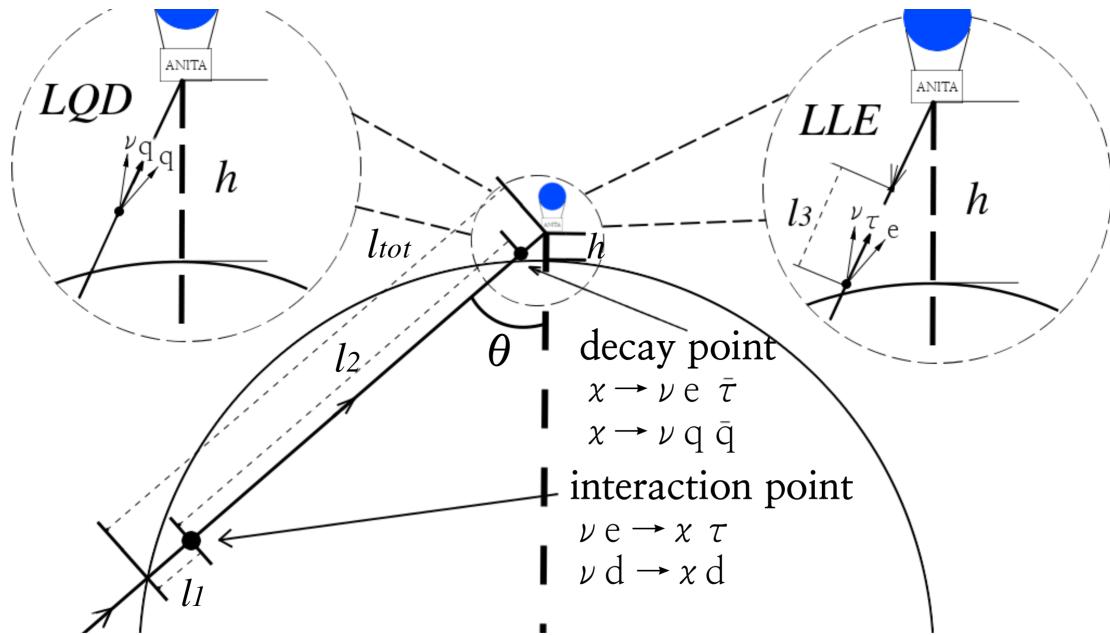


M. Carena, D. Choudhury, S. Lola, C. Quigg, Hep-ph/9804380;
P. S. Bhupal Dev, Dilip Kumar Ghosh, Werner Rodejohann.1605.09743

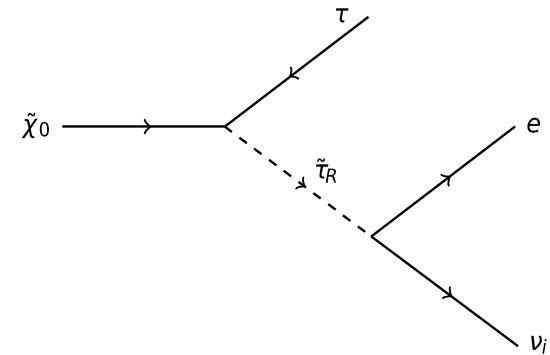
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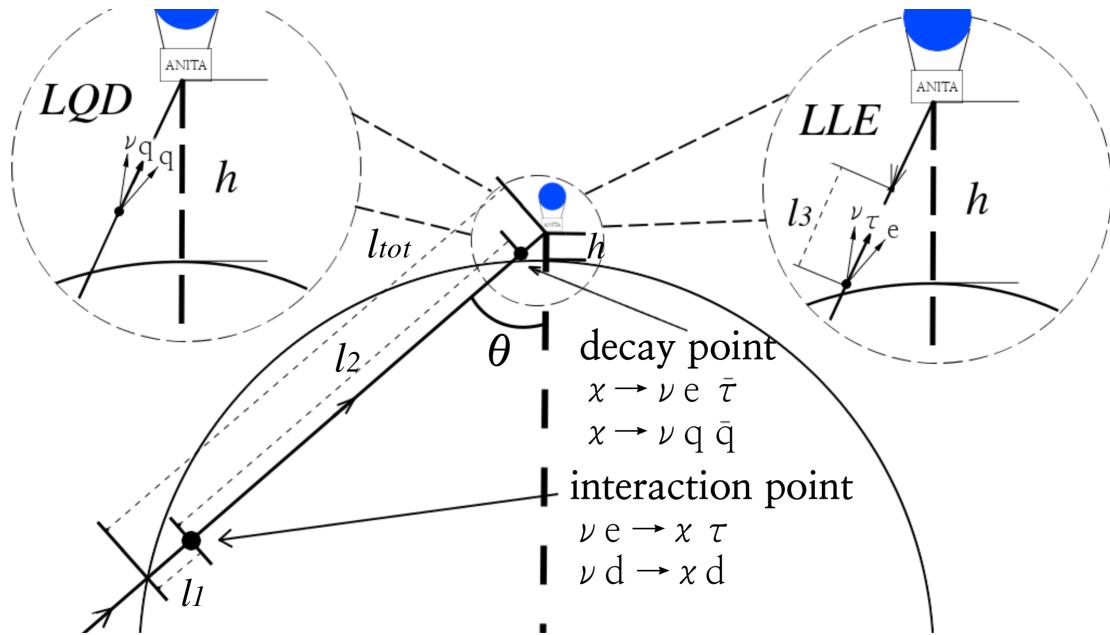
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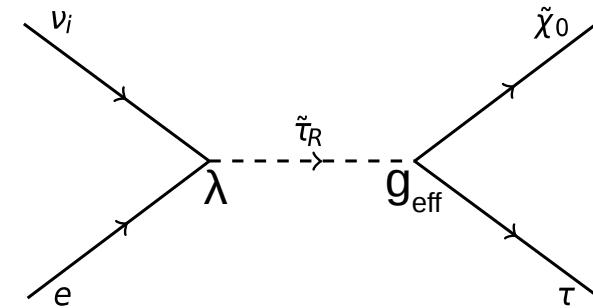
New Physics Interpretation to ANITA-anomaly

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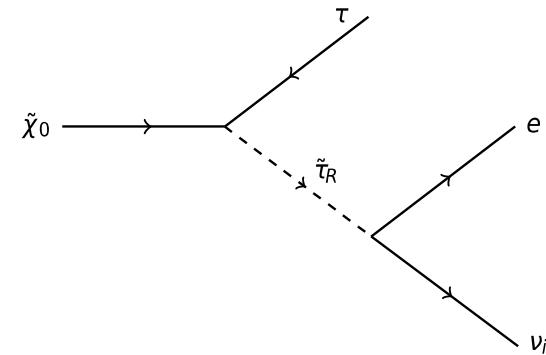
Jack Collins and Bhupal Dev, Yicong Sui, 1810.08479



RPV-SUSY



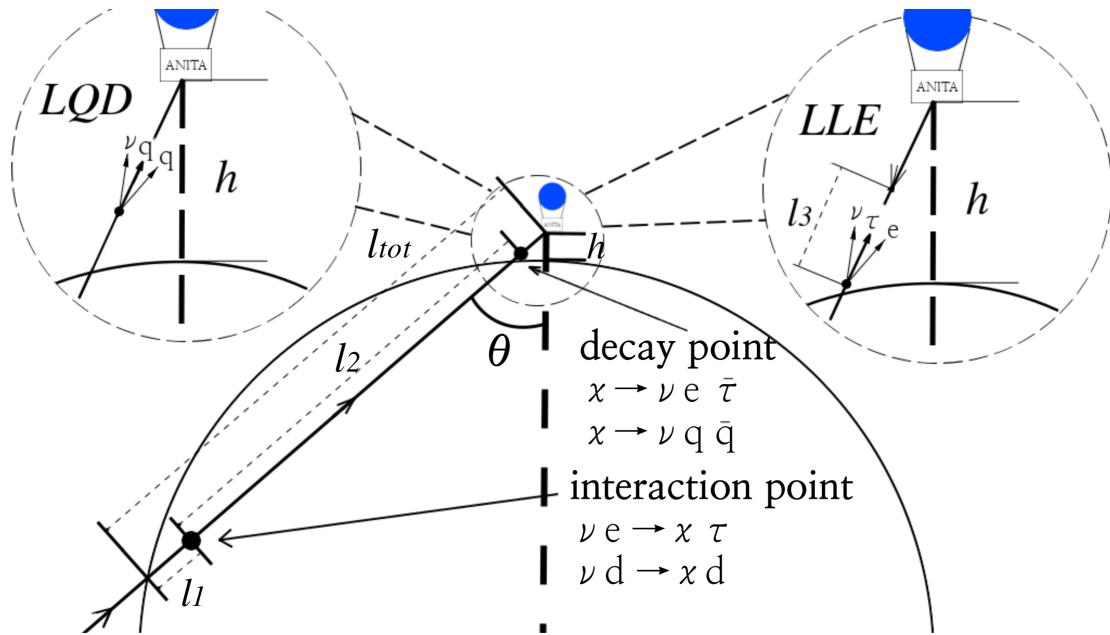
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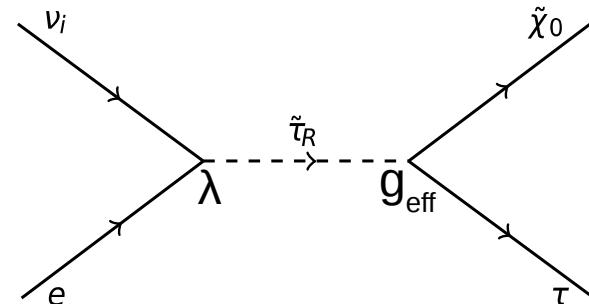
New Physics Interpretation to ANITA-anomaly

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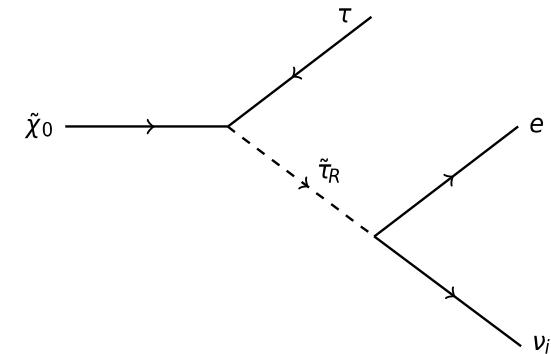
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RPV-SUSY



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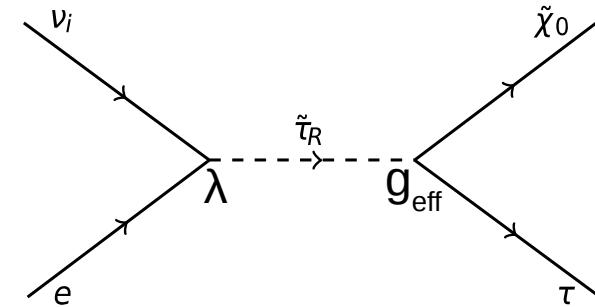
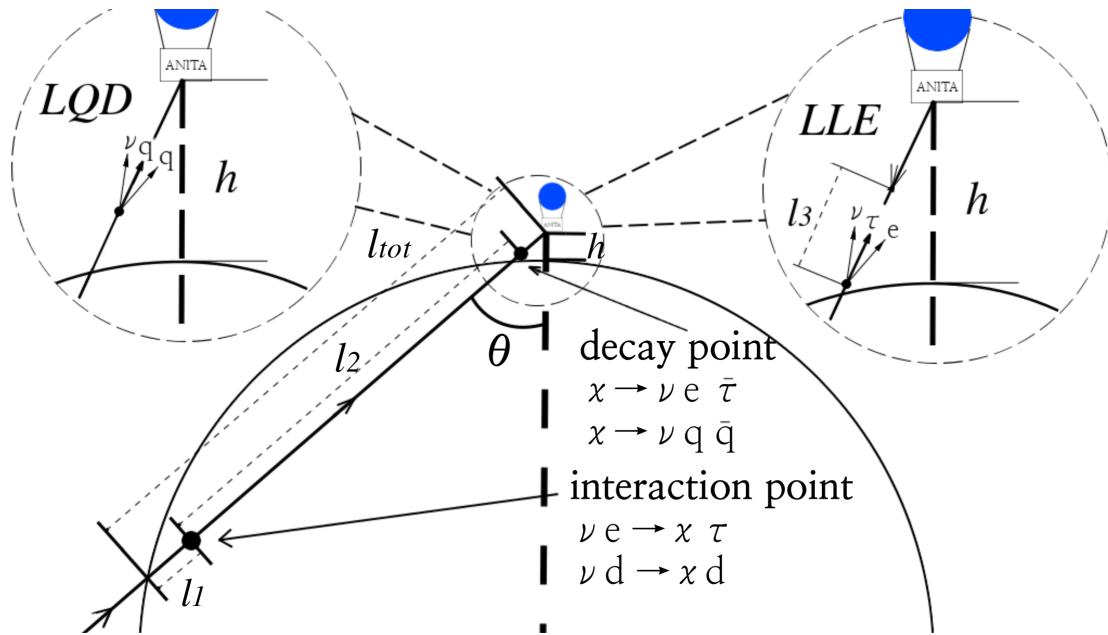
$$W_{RPV} = \lambda_{ijk} L^i L^j \bar{E}^k + \lambda'_{ijk} L^i Q^j \bar{D}^k + \lambda''_{ijk} \bar{U}^i \bar{D}^j \bar{D}^k$$



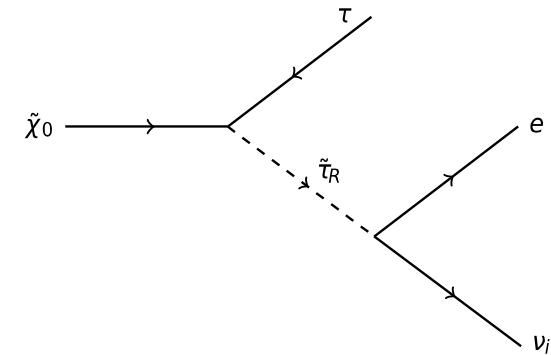
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RPV-SUSY

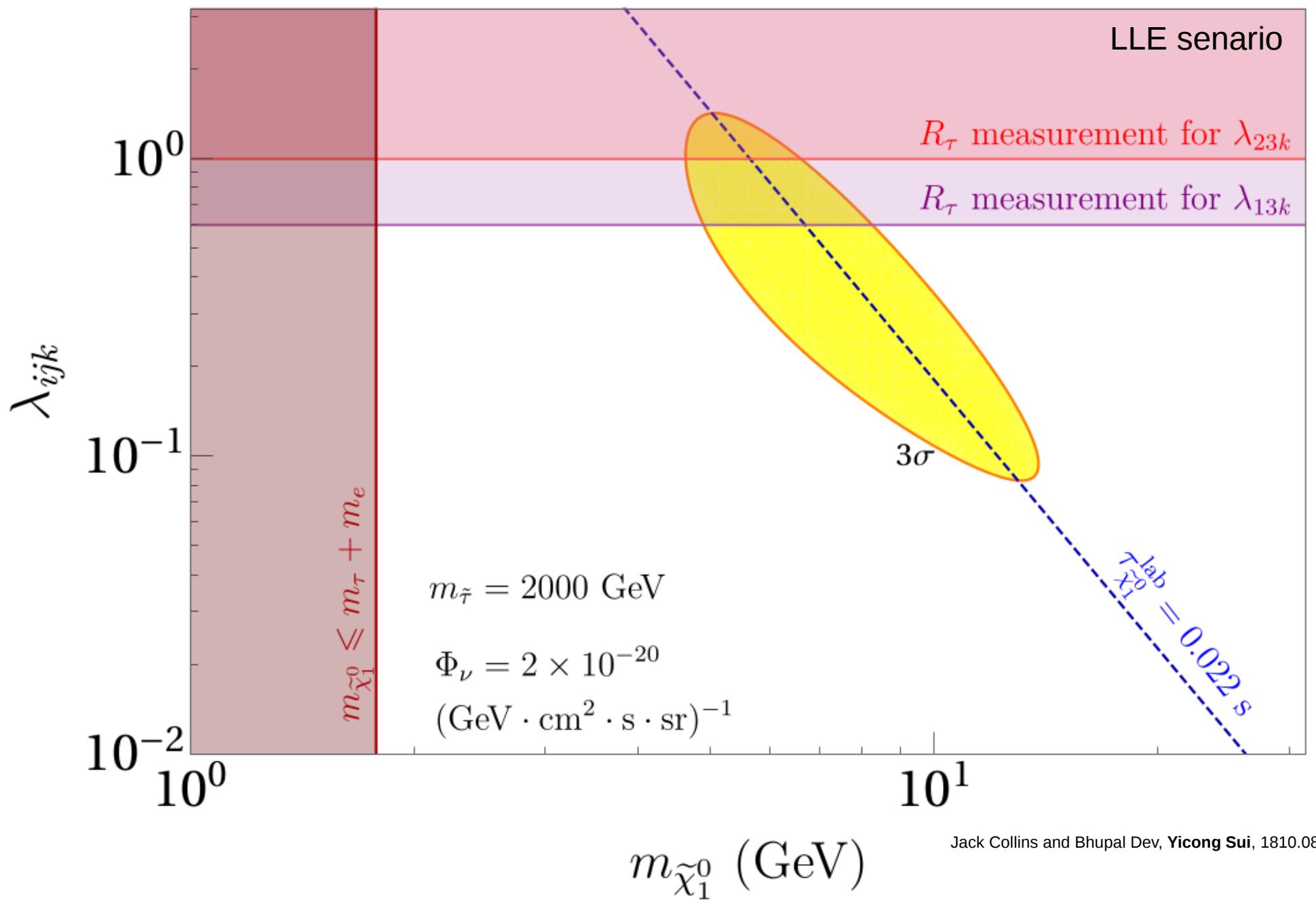
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$$\begin{aligned} \sigma_{RPV} &= \frac{8\pi}{M_{\tilde{\tau}}^2} \text{Br}[\tilde{\tau} \rightarrow \nu + e] \cdot \text{Br}[\tilde{\tau} \rightarrow \chi + \tau] \\ &= \frac{8\pi}{M_{\tilde{\tau}}^2} \frac{|\lambda|^2}{|\lambda|^2 + g_{eff}^2} \frac{g_{eff}^2}{|\lambda|^2 + g_{eff}^2} \end{aligned}$$

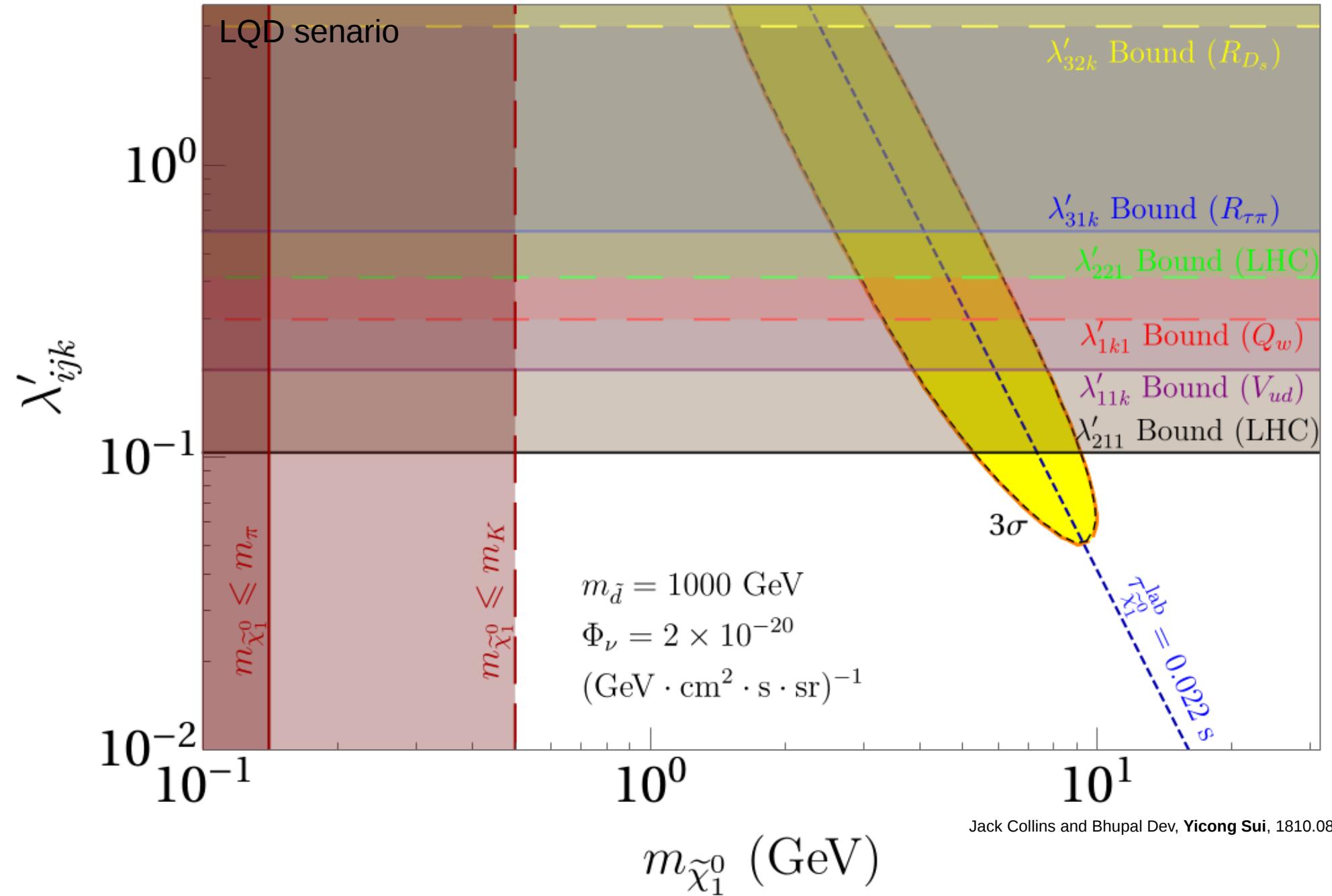
$$\Gamma(\chi \rightarrow \tau e \nu) \sim \frac{3\alpha \lambda_{i31}^2}{128\pi^2} \frac{M_\chi^5}{M_{\tilde{\tau}}^4}$$



Washington University in St. Louis



Jack Collins and Bhupal Dev, Yicong Sui, 1810.08479



RPV-SUSY Interpretation to B-anomaly: $R_D R_{D^*}$

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$$\mathcal{L}(b \rightarrow c \ell_l \bar{\nu}_{l'}) = -\frac{4G_F}{\sqrt{2}} V_{cb} (\delta_{ll'} + \Delta_{ll'}^c) \bar{\ell}_L^\ell \gamma^\mu \nu_L^{l'} \bar{c}_L \gamma^\mu b_L \quad \Delta_{ll'}^c = \sum_{j=1}^3 \frac{\sqrt{2}}{4G_F} \frac{\lambda'_{l'33} \lambda'_{lj3}}{2m_{\tilde{b}_R}^2} \frac{V_{cj}}{V_{cb}}$$

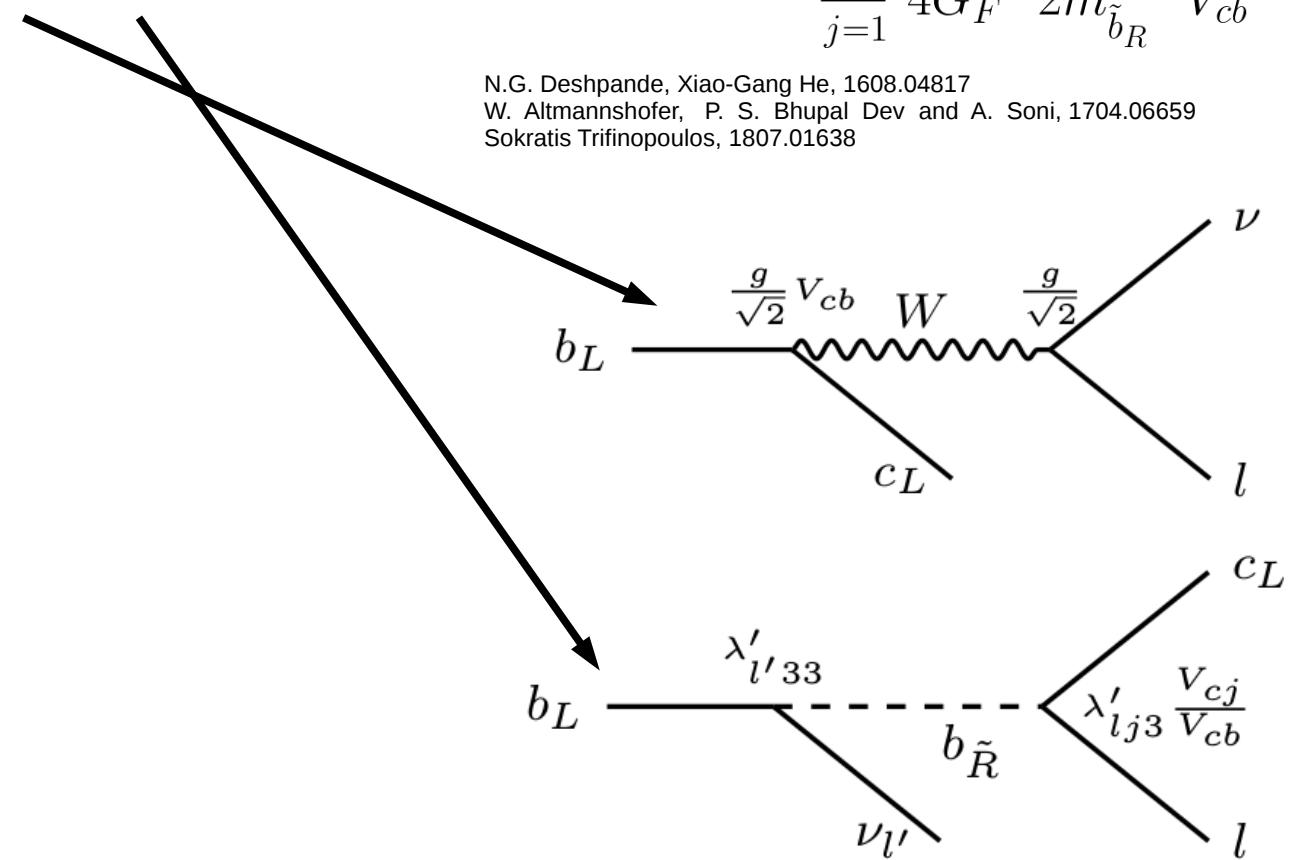
N.G. Deshpande, Xiao-Gang He, 1608.04817
W. Altmannshofer, P. S. Bhupal Dev and A. Soni, 1704.06659
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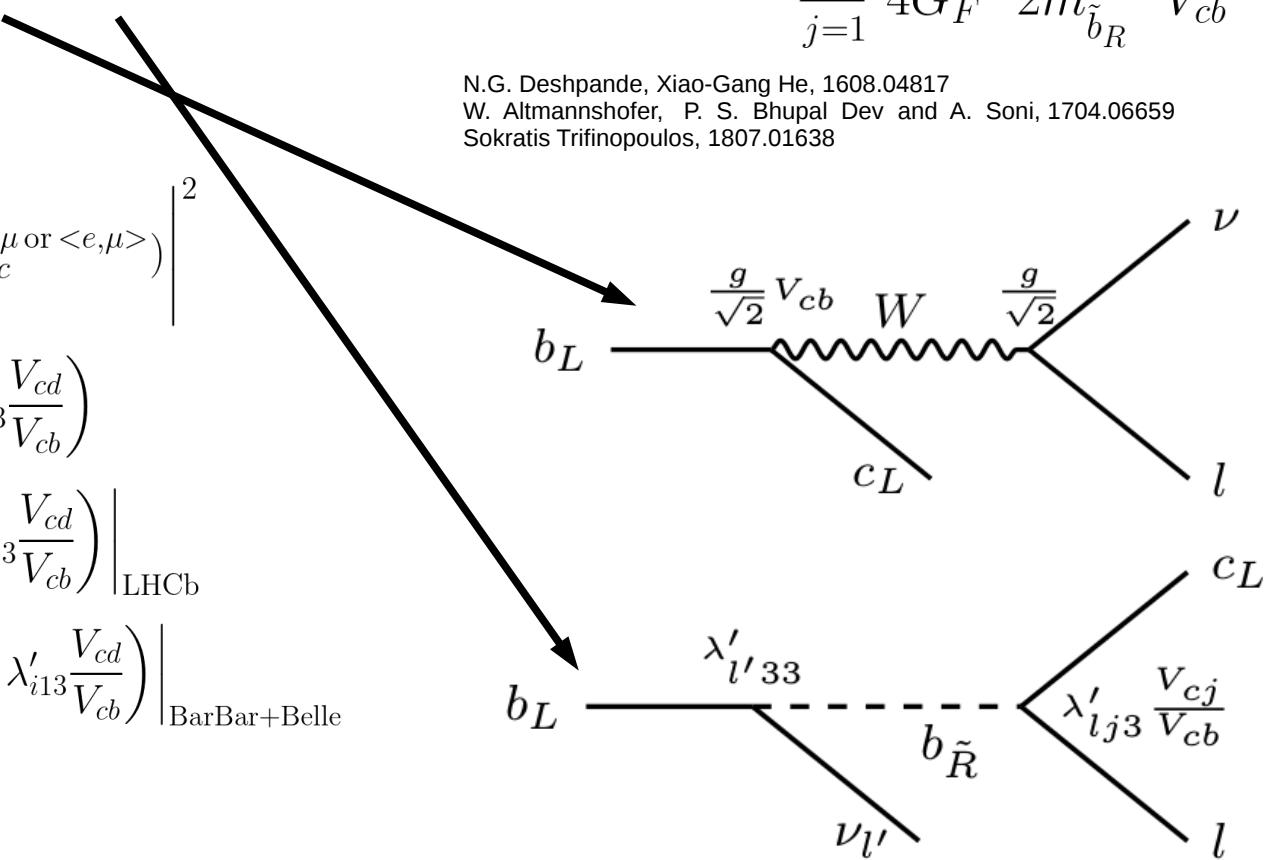
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$$\frac{R_D}{R_D^{\text{SM}}} = \frac{R_{D^*}}{R_{D^*}^{\text{SM}}} = \left| 1 + \frac{v^2}{2m_{\tilde{b}_R}^2} \text{Re}(X_c^\tau - X_c^{\mu \text{ or } <e,\mu>}) \right|^2$$

$$X_c^\tau = \left(\sum_{i=1}^3 \lambda'_{i33} \right) \left(\lambda'_{333} + \lambda'_{323} \frac{V_{cs}}{V_{cb}} + \lambda'_{313} \frac{V_{cd}}{V_{cb}} \right)$$

$$X_c^\mu = \left(\sum_{i=1}^3 \lambda'_{i33} \right) \left(\lambda'_{233} + \lambda'_{223} \frac{V_{cs}}{V_{cb}} + \lambda'_{213} \frac{V_{cd}}{V_{cb}} \right) \Big|_{\text{LHCb}}$$

$$X_c^{<e,\mu>} = \left(\sum_{i=1}^3 \lambda'_{i33} \right) \frac{1}{2} \sum_{i=1}^2 \left(\lambda'_{i33} + \lambda'_{i23} \frac{V_{cs}}{V_{cb}} + \lambda'_{i13} \frac{V_{cd}}{V_{cb}} \right) \Big|_{\text{BarBar+Belle}}$$



RPV-SUSY Interpretation to B-anomaly: $R_K R_{K^*}$

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Diganta Das, Chandan Hati, Girish Kumar, and Namit Mahajan, 1705.09188
Kevin Earla, Thomas Gr'egoirea, 1806.01343

$$\mathcal{H}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{4\pi} V_{tb} V_{ts}^* \sum_{i=9,10} (C_i \mathcal{O}_i + C'_i \mathcal{O}'_i)$$

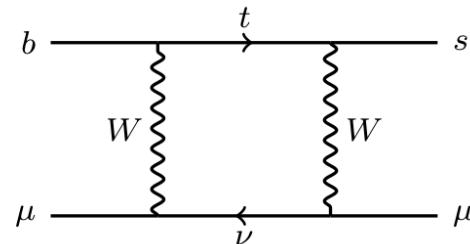
$$\begin{aligned}\mathcal{O}_{9(10)} &= (\bar{s}_L \gamma^\mu b_L)(\bar{\ell} \gamma_\mu (\gamma_5) \ell), \\ \mathcal{O}'_{9(10)} &= (\bar{s}_R \gamma^\mu b_R)(\bar{\ell} \gamma_\mu (\gamma_5) \ell)\end{aligned}$$

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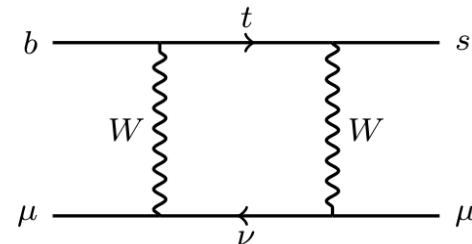
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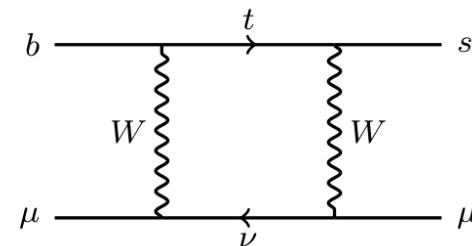
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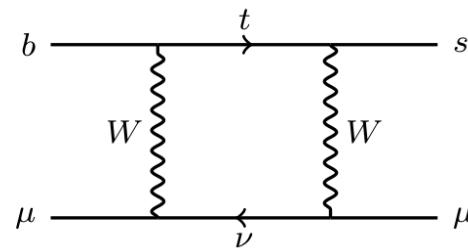
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$$\begin{aligned}(\delta C_9)^\mu = -(\delta C_{10})^\mu &= \frac{m_t^2}{m_{\tilde{b}_R}^2} \frac{|\lambda'_{233}|^2}{16\pi\alpha_{\text{em}}} - \frac{v^2}{16m_{\tilde{b}_R}^2} \frac{X_{bs} X_{\mu\mu}}{e^2 V_{tb} V_{ts}^*} \\ &- \frac{v^2}{16(m_{\tilde{t}_L}^2 - m_{\tilde{\nu}_\tau}^2)} \log \left(\frac{m_{\tilde{t}_L}^2}{m_{\tilde{\nu}_\tau}^2} \right) \frac{X_{bs} X_{\mu\mu}}{e^2 V_{tb} V_{ts}^*}\end{aligned}$$



$$X_{bs} = \sum_{i=1}^3 \lambda'_{i33} \lambda'_{i23}, \quad X_{\mu\mu} = \sum_{j=1}^3 |\lambda'_{2j3}|^2$$

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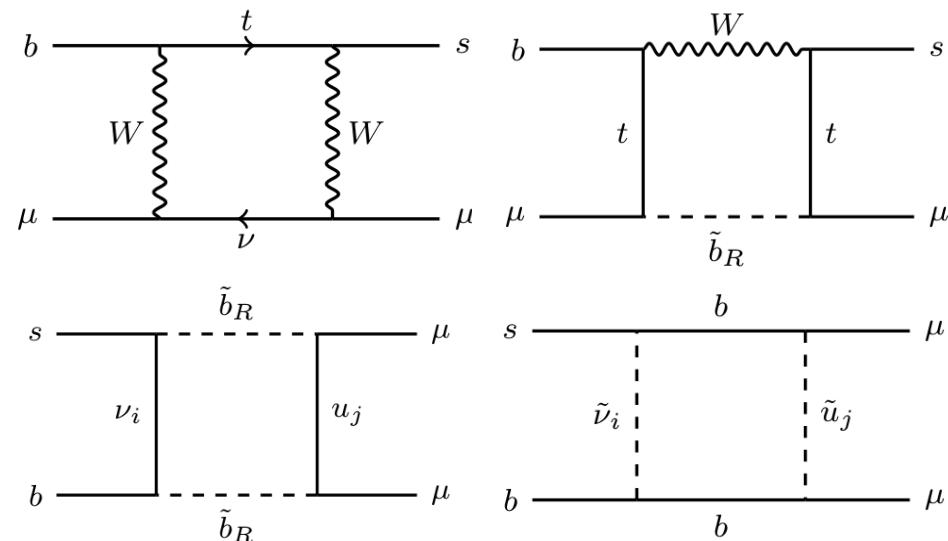
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RPV-SUSY Interpretation to B-anomaly: $R_K R_{K^*}$

Diganta Das, Chandan Hati, Girish Kumar, and Namit Mahajan, 1705.09188
 Kevin Earla, Thomas Grégoirea, 1806.01343

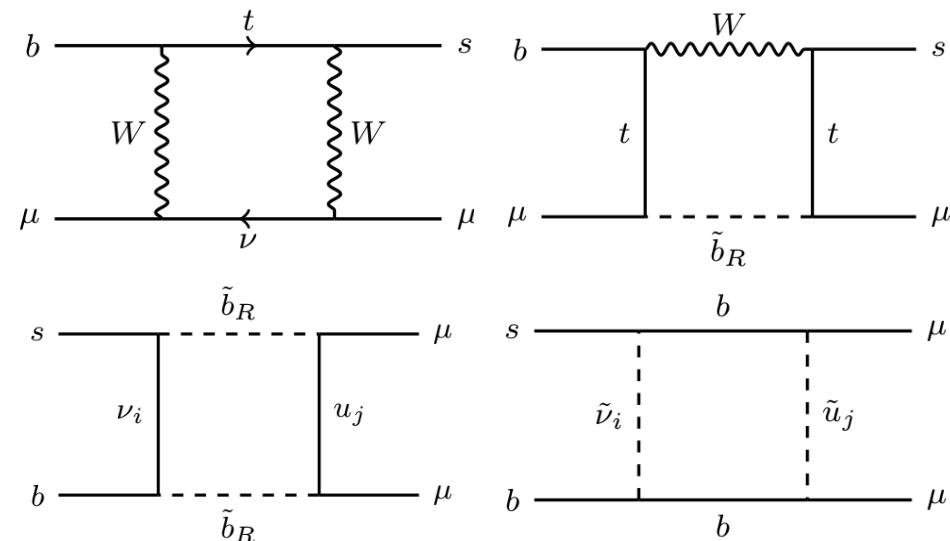
$$\mathcal{H}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{4\pi} V_{tb} V_{ts}^* \sum_{i=9,10} (C_i \mathcal{O}_i + C'_i \mathcal{O}'_i)$$

$$\begin{aligned}\mathcal{O}_{9(10)} &= (\bar{s}_L \gamma^\mu b_L)(\bar{\ell} \gamma_\mu (\gamma_5) \ell), \\ \mathcal{O}'_{9(10)} &= (\bar{s}_R \gamma^\mu b_R)(\bar{\ell} \gamma_\mu (\gamma_5) \ell)\end{aligned}$$

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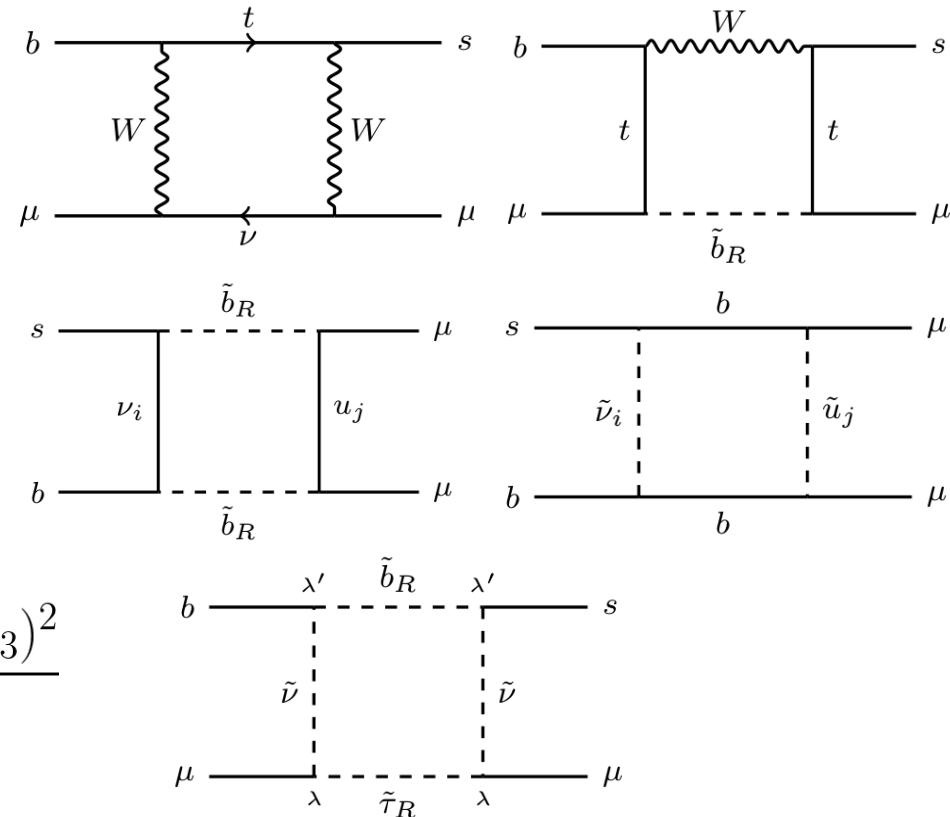
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$$\text{Fourth Term} = -\frac{v^2 \log\left(\frac{m_{\tilde{b}_R}^2}{m_{\tilde{\tau}_R}^2}\right)}{8(m_{\tilde{b}_R}^2 - m_{\tilde{\tau}_R}^2)} \frac{\lambda'_{323} \lambda'_{333} (\lambda_{323})^2}{e^2 V_{tb} V_{ts}^*}$$



Sokratis Trifinopoulos, 1807.01638



Anita, Anomaly



B, Anomaly

Hey, B. Are we
related?





Anita, Anomaly



B, Anomaly

Hey, B. Are we
related?





Washington University in St. Louis

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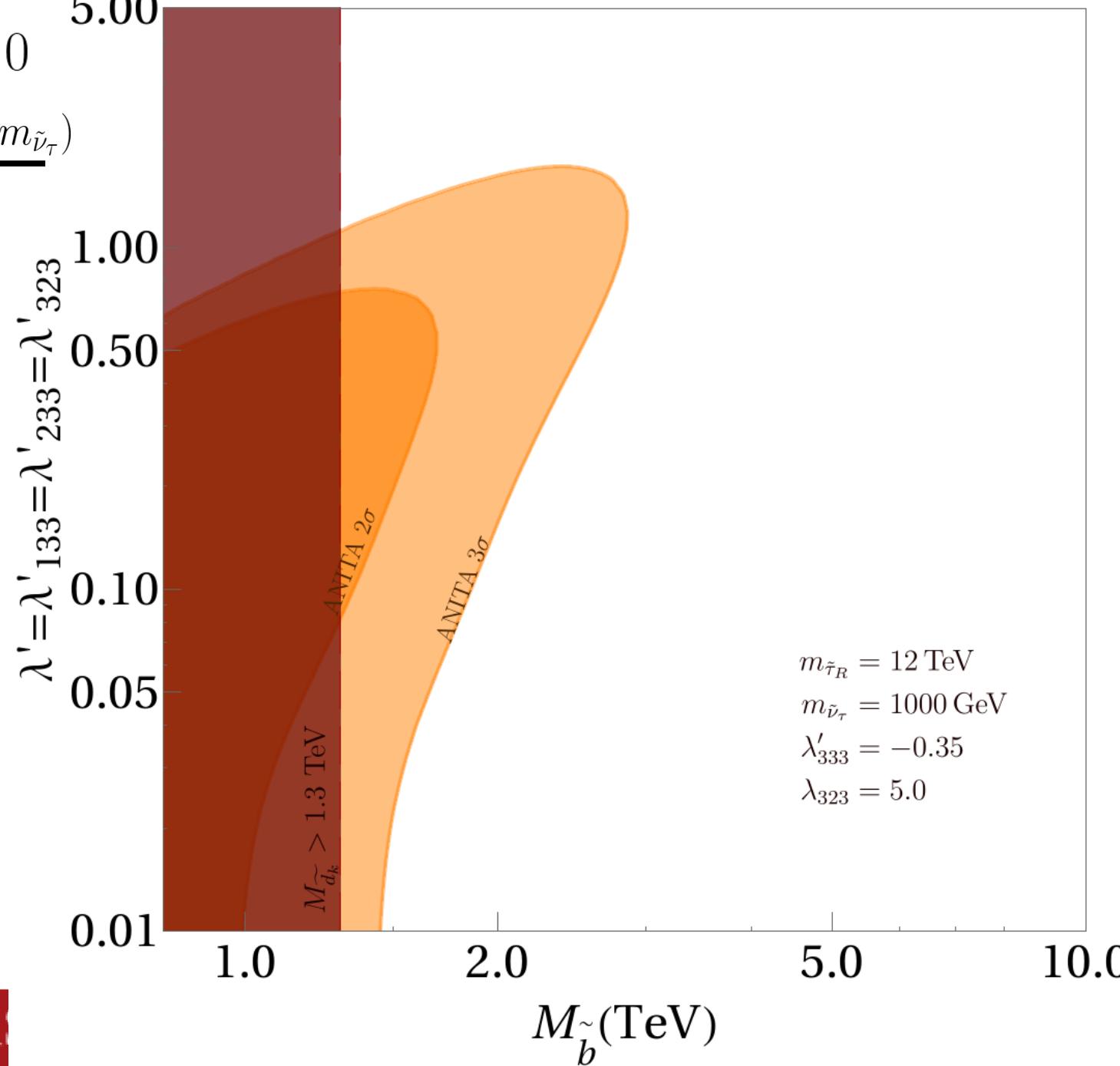
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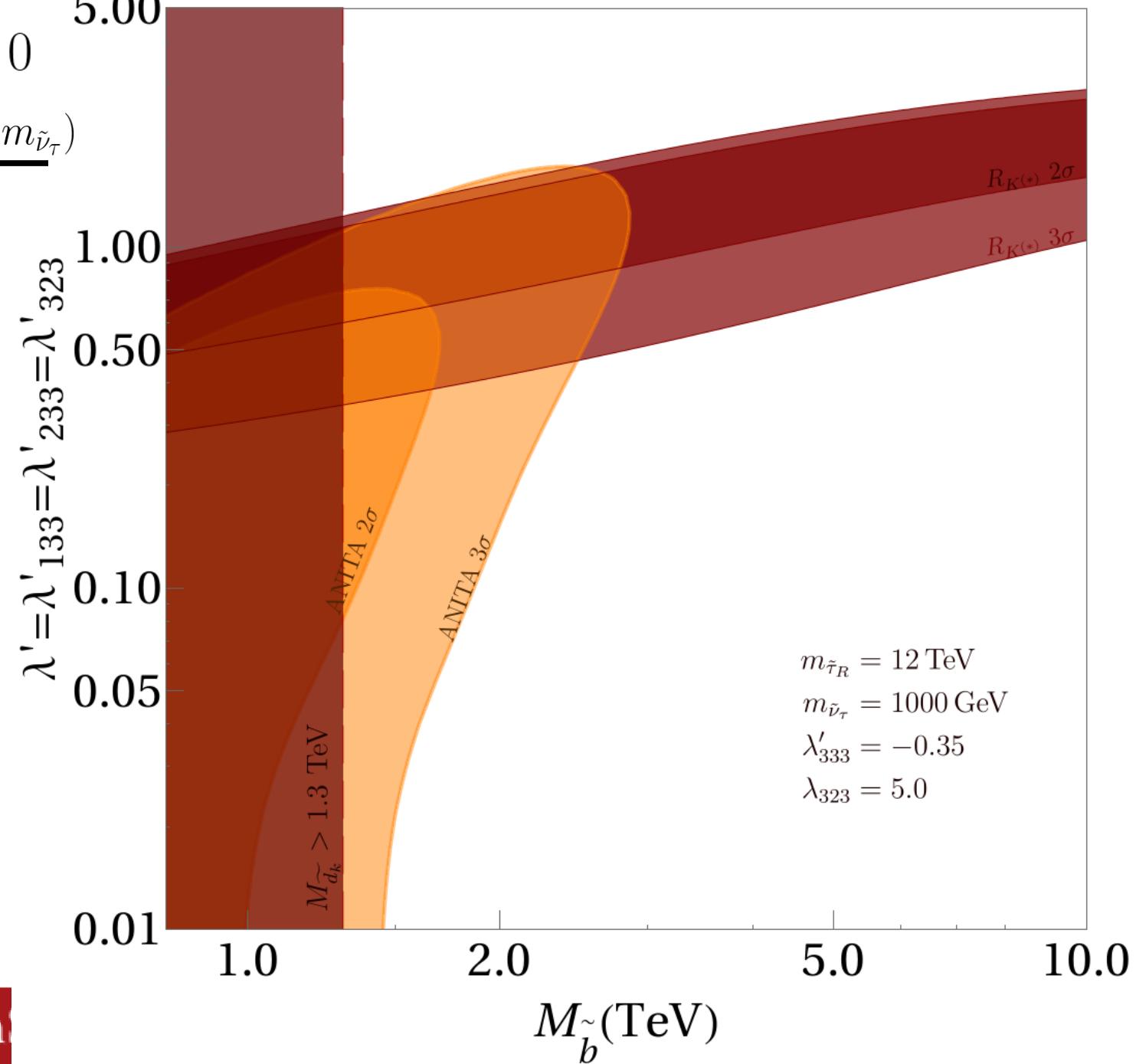
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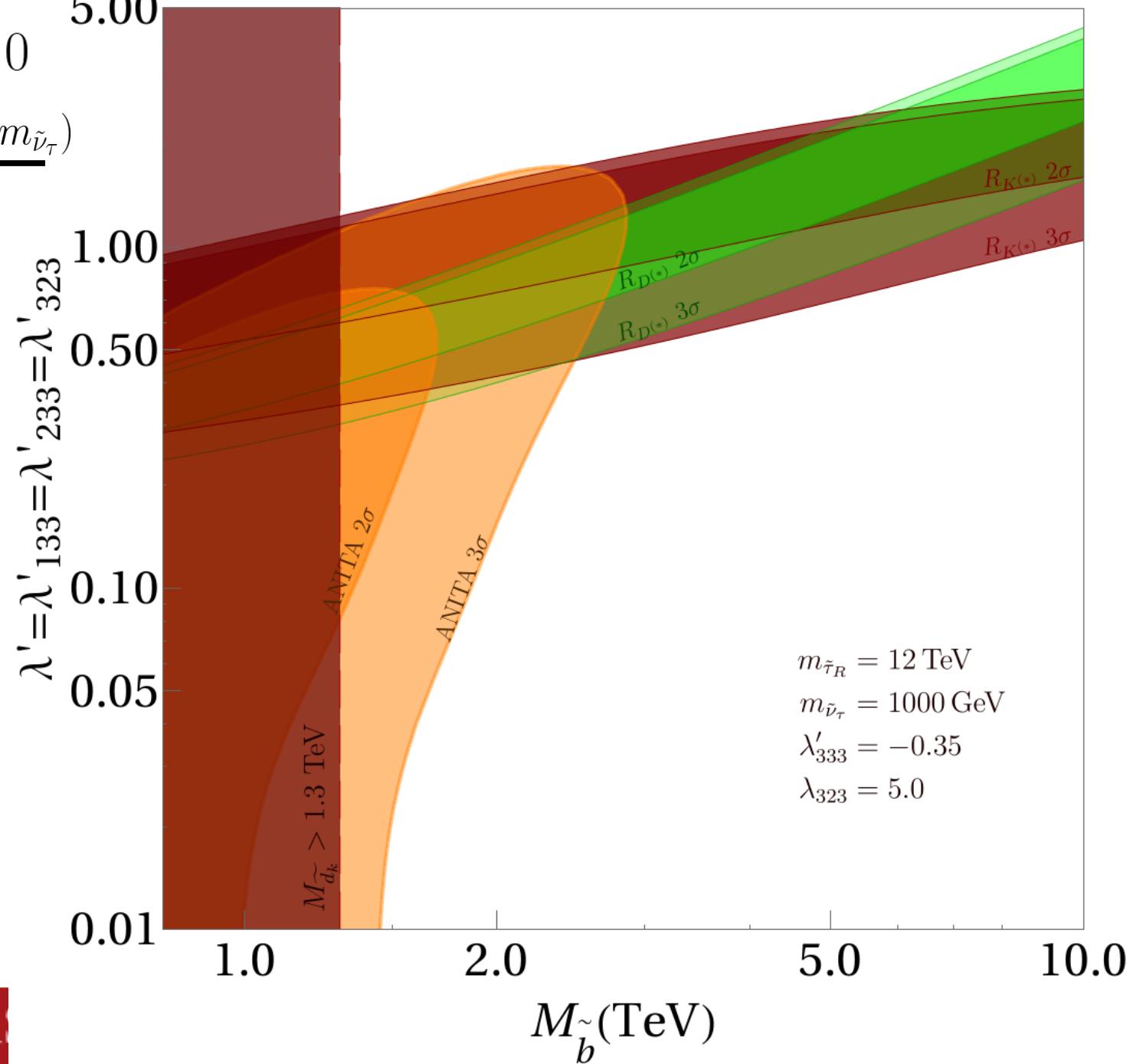
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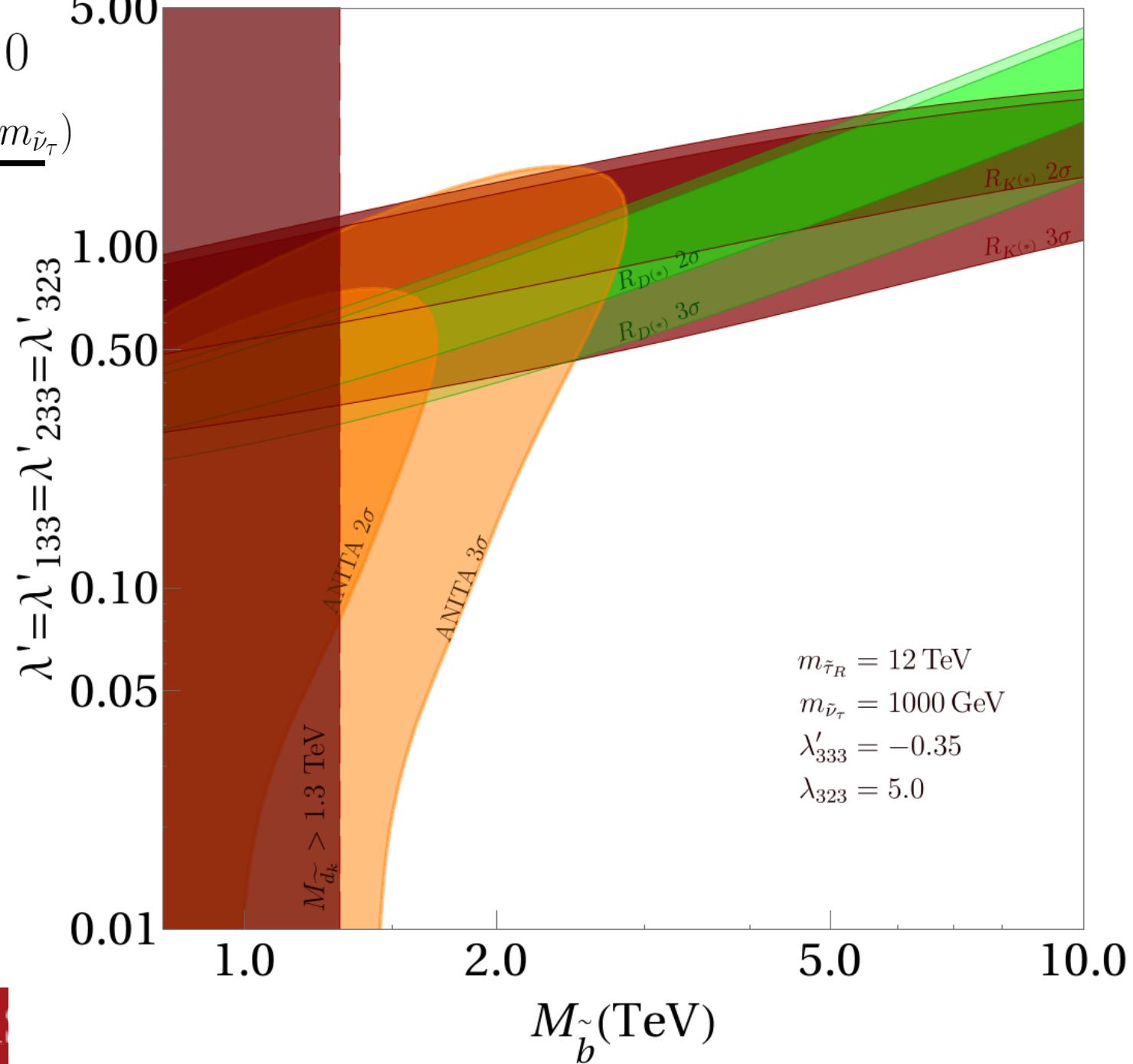
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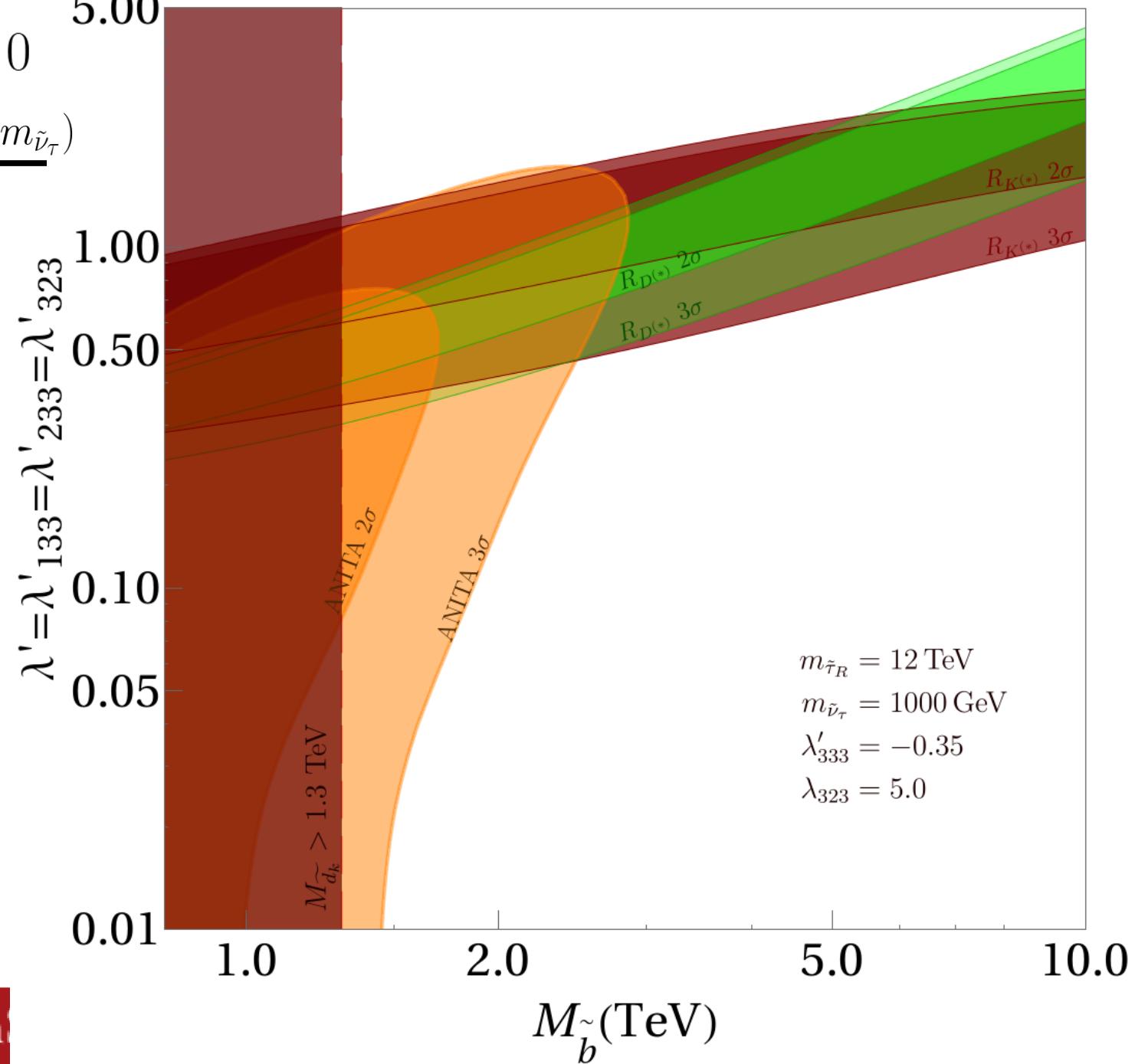
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Sokratis Trifinopoulos,
1807.01638



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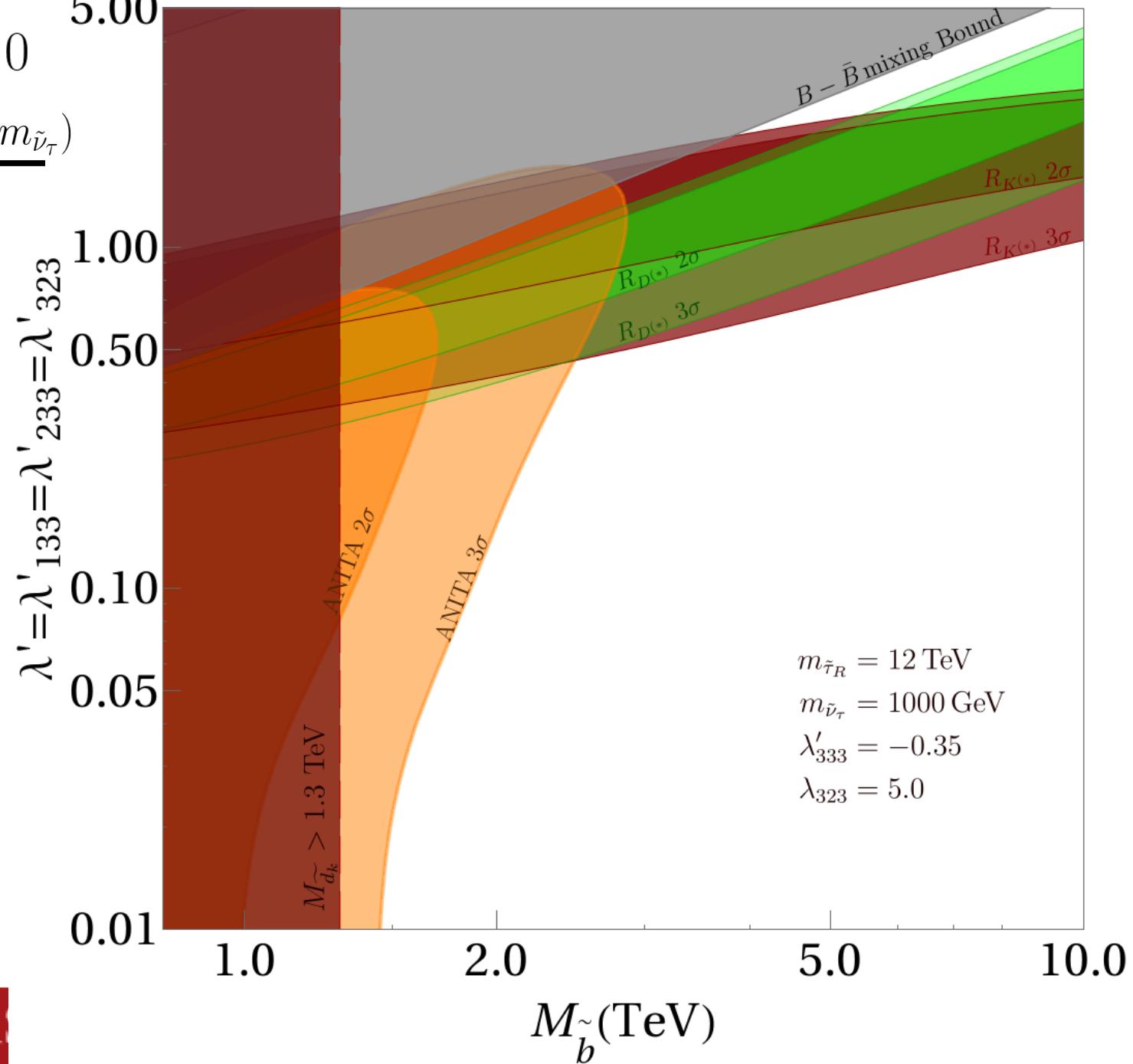
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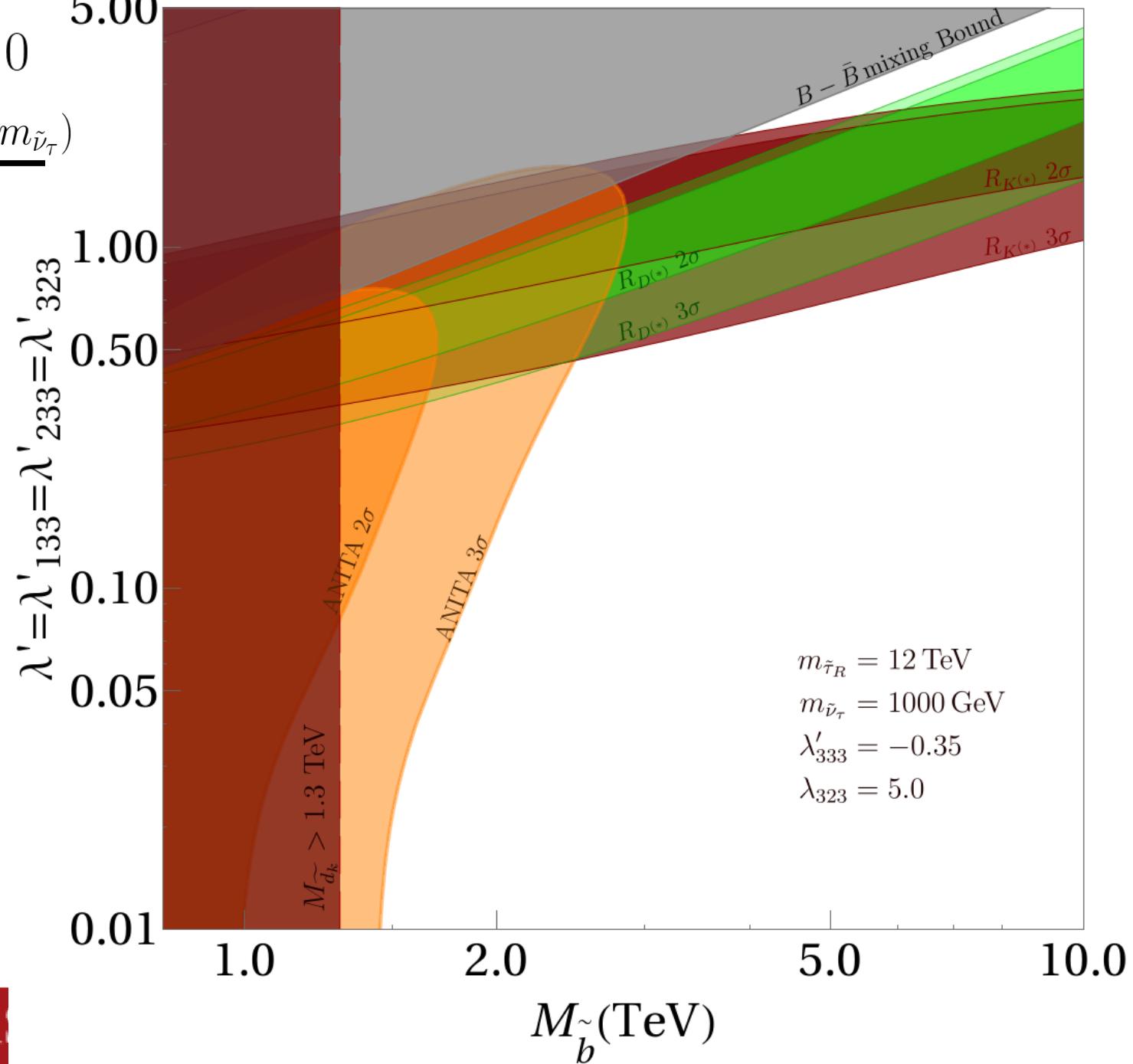
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Dario Buttazzo, et al,
1706.07808



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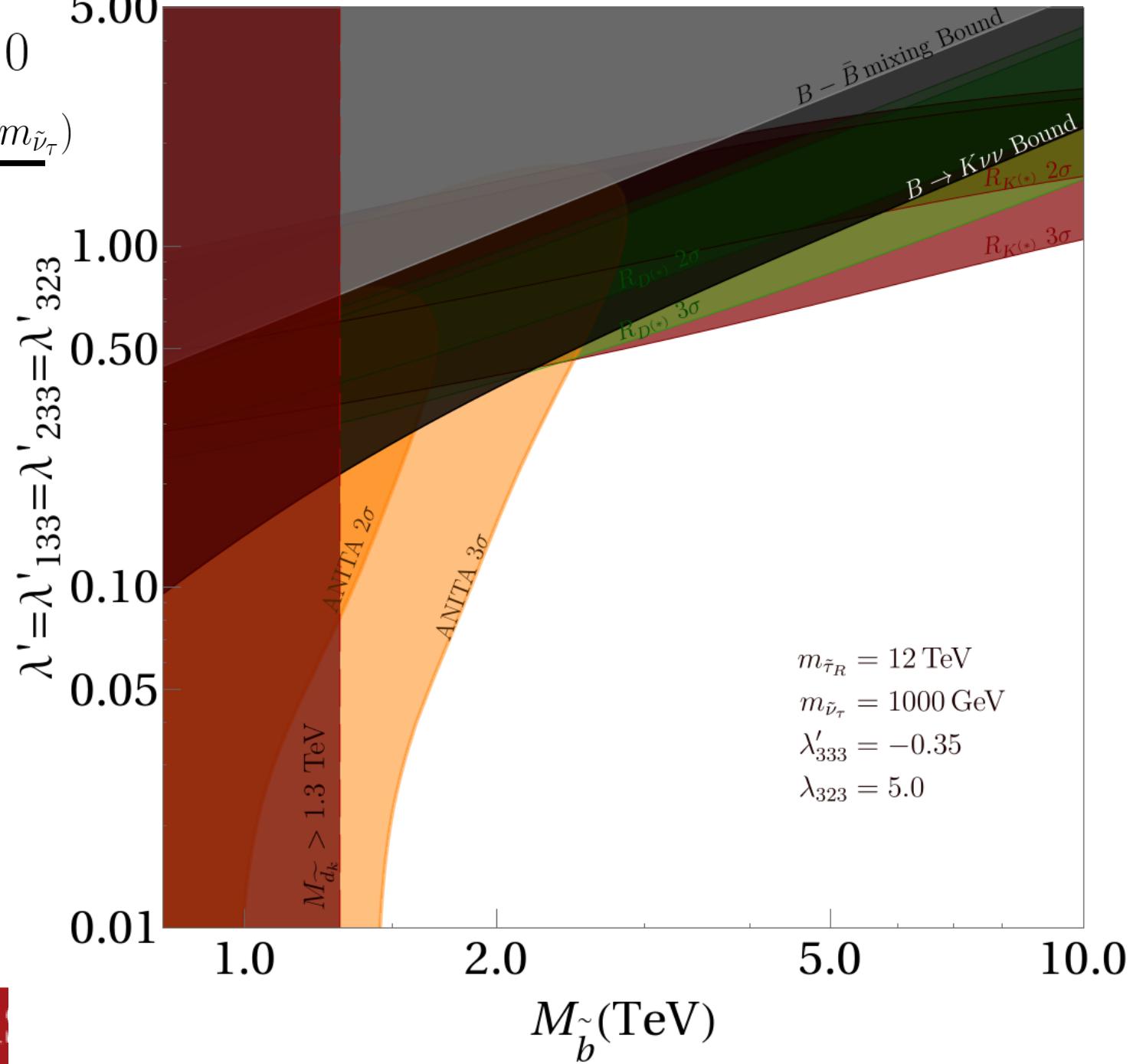
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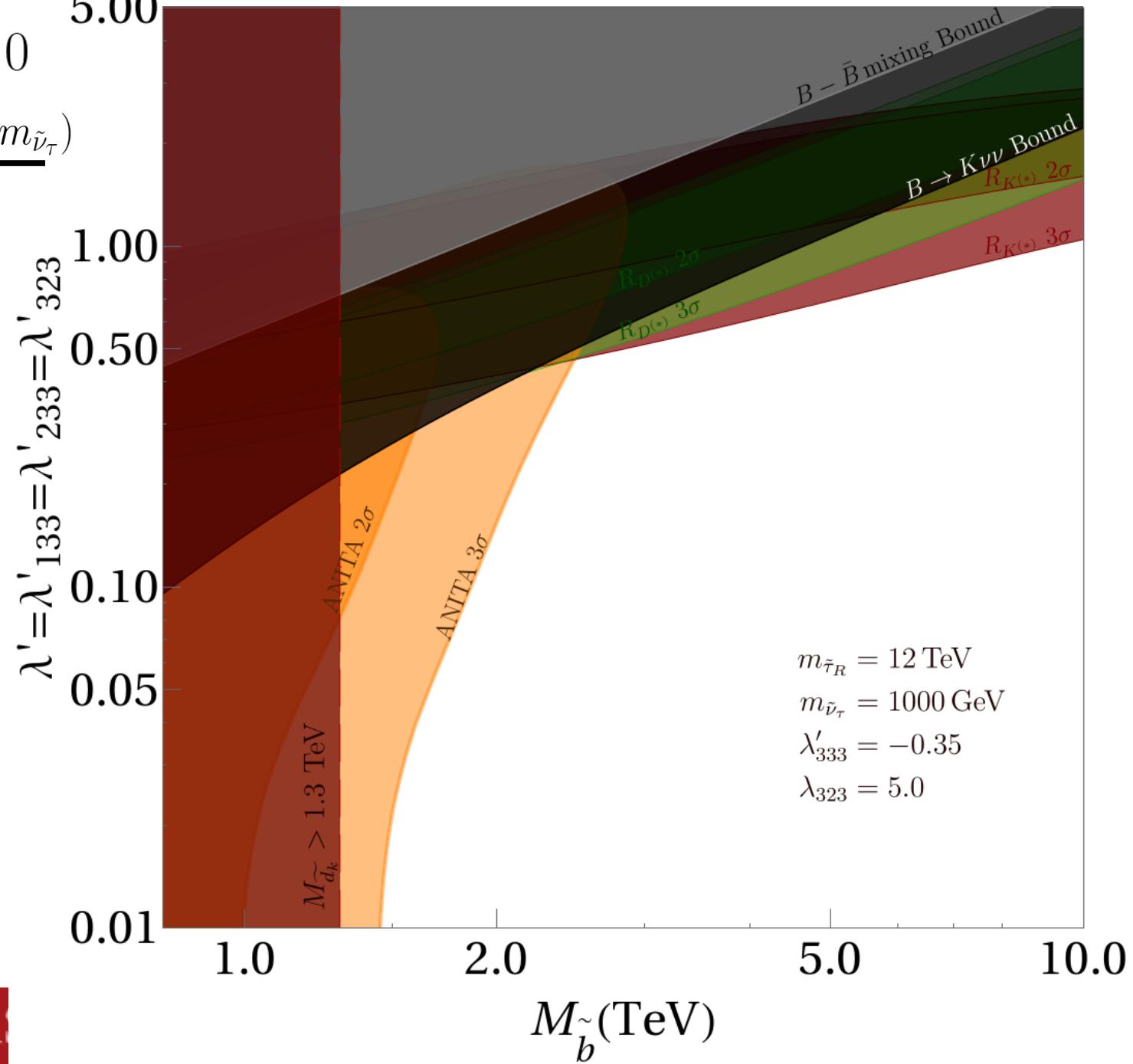
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Heavy Flavor
Averaging Group, et al,
1612.07233



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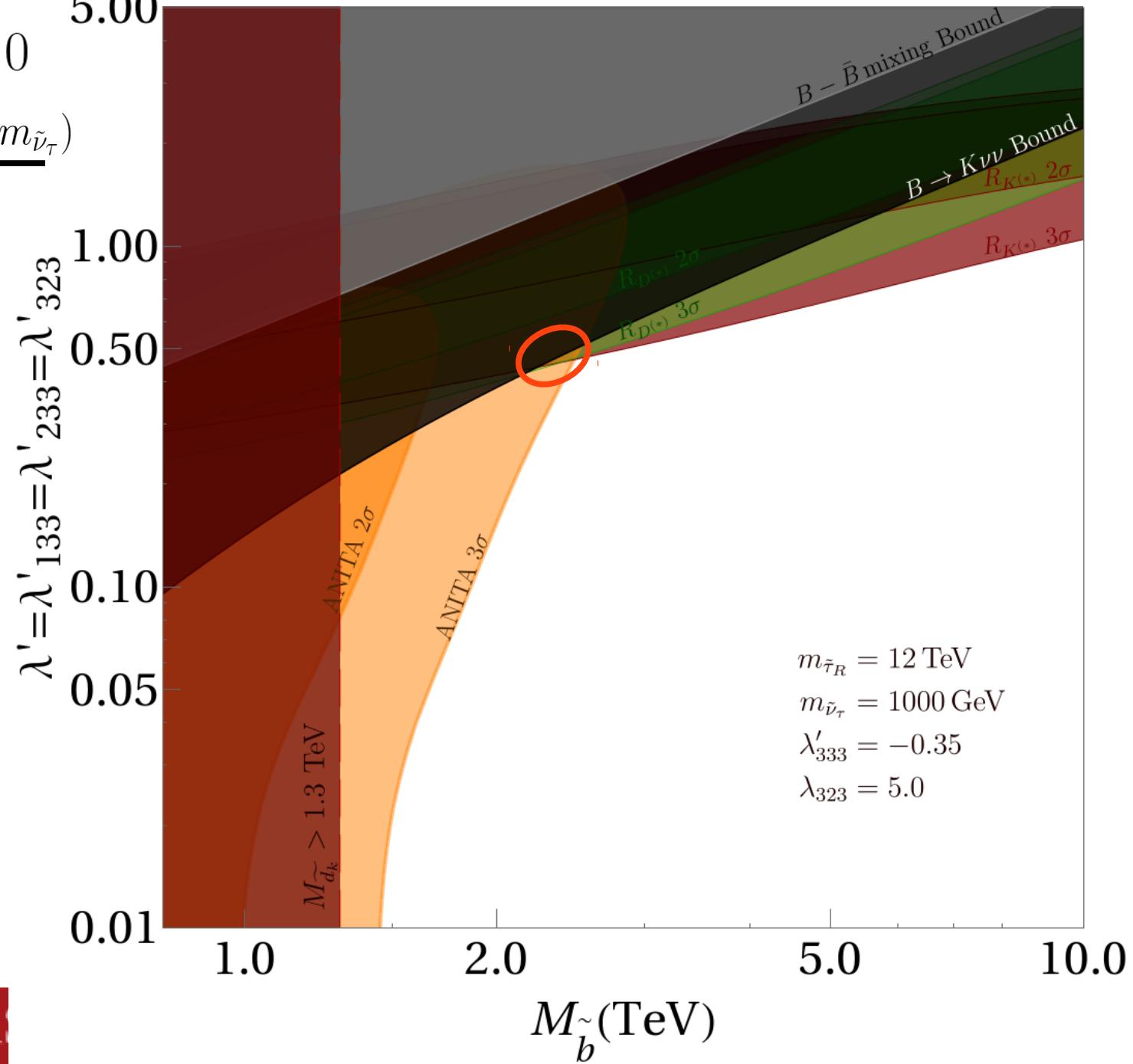
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Conclusion:

- 1.Under the framework of RPV-SUSY, ANITA anomaly has quite large parameter space, which is in the similar range demanded by RD and RK.
- 2.Rk Rk* and RD RD* anomalies could be explained simultaneously after a “fourth term” is included in the traditional RPV-SUSY Rk Rk* expression.
- 3.Under the simplified parameter setup, we find that there exist parameter spaces that could satisfy RK-RD-ANITA altogether.
- 4.Relaxing the parameter setup and letting more lambda' running free could possibly lead to larger preferred region for the parameters.
- 5.This framework could also be expanded to include muon g-2 anomaly

Thank you



BK slides-Constraints

BK slides-Constraints

$$R_{B \rightarrow K^{(*)} \nu \bar{\nu}} = \sum_{i=i'=1}^3 \frac{1}{3} \left| 1 + \frac{\Delta_{\nu_i \bar{\nu}'_i}^{\text{RPV}}}{X t V_{ts}^* V_{tb}} \right|^2 + \sum_{i \neq i'} \frac{1}{3} \left| \frac{\Delta_{\nu_i \bar{\nu}'_i}^{\text{RPV}}}{X t V_{ts}^* V_{tb}} \right|^2$$
$$\Delta_{\nu_i \bar{\nu}'_i}^{\text{RPV}} = \frac{\pi s_W^2}{\sqrt{2} G_F \alpha} \left| -\frac{\lambda'_{i33} \lambda'_{i'23}}{2 m_{\tilde{b}_R}^2} \right|$$

BK slides-Constraints

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BK slides-Constraints

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BK slides-Constraints

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$B - \bar{B}$ mixing

$$\Delta M_B = \frac{2 m_B F_B^2}{3} \left| P_1^{VLL} \frac{\lambda'_{i23} \lambda'_{i33} \lambda'_{j33} \lambda'_{j23}}{128 \pi^2 m_{\tilde{b}_R}^2} \right|$$

BK slides-Constraints

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