

NSCL 8/02/05

Spectroscopic factors and the physics of the single-particle strength distribution in nuclei

- Lecture 1: 7/18/05 Propagator description of single-particle motion and the link with experimental data
- Lecture 2: 7/19/05 From diagrams to Hartree-Fock and spectroscopic factors < 1
- Lecture 3: 7/20/05 Influence of long-range correlations and the relation to excited states
- Lecture 4: 7/27/05 Role of short-range and tensor correlations associated with realistic interactions.
- Lecture 5: 7/28/05 Summary of results. Prospects for nuclei with N very different from Z .
- Lecture 6: 8/02/05 Saturation problem of nuclear matter

Wim Dickhoff
Washington University in St. Louis

The two “most elusive” numbers in nuclear physics

- What are these numbers?
- In what sense are they elusive?
- What is the history?
- Three-body forces? Relativity? Give up?
- What has been learned from $(e,e'p)$?
- What really decides the saturation density?
- Nuclear Matter with SRC? No LRC?
- Conclusions

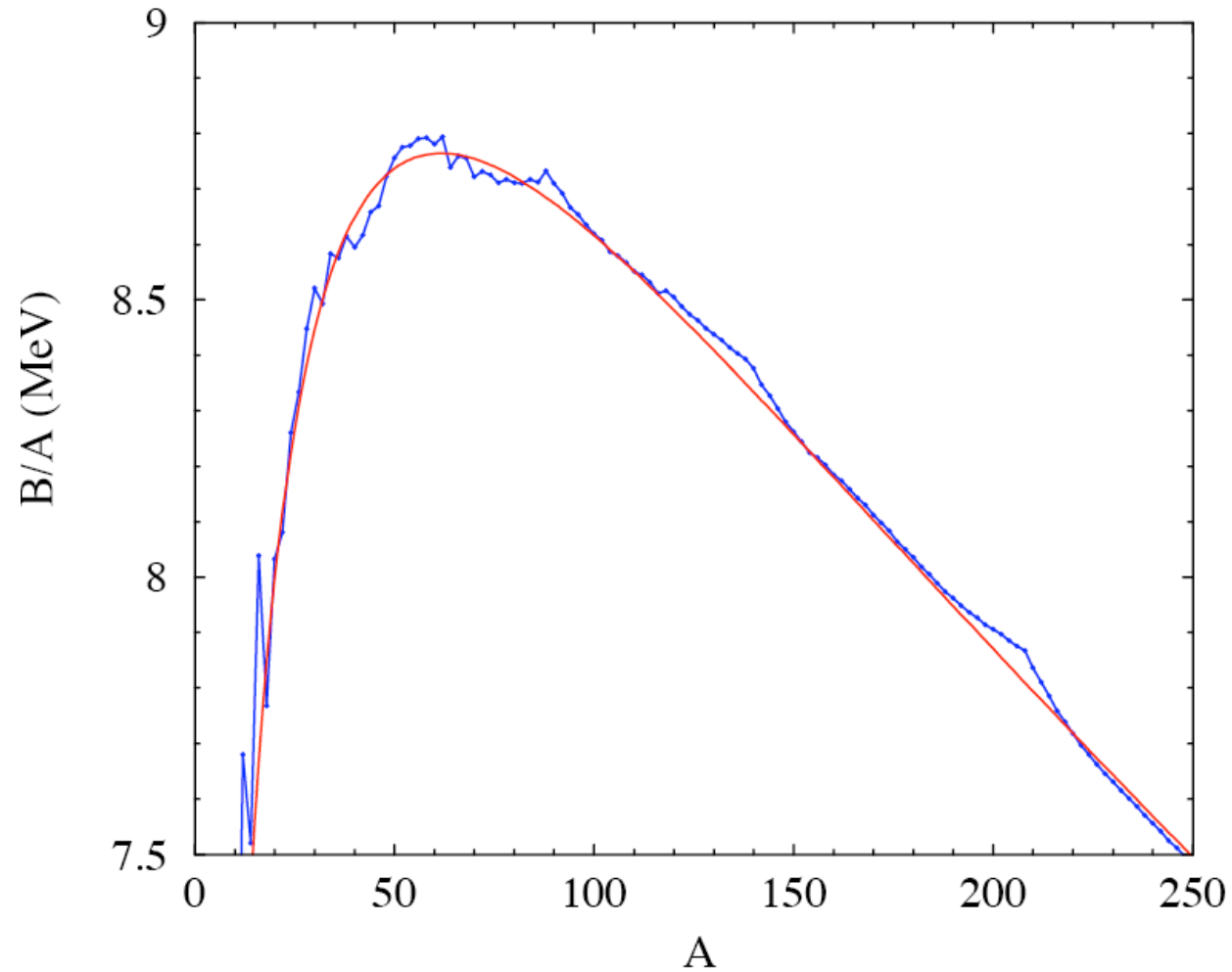
Empirical Mass Formula

Global representation of nuclear masses (Bohr & Mottelson)

$$B = b_{vol}A - b_{surf}A^{2/3} - \frac{1}{2}b_{sym} \frac{(N - Z)^2}{A} - \frac{3}{5} \frac{Z^2 e^2}{R_c}$$

- Volume term $b_{vol} = 15.56 \text{ MeV}$
- Surface term $b_{surf} = 17.23 \text{ MeV}$
- Symmetry energy $b_{sym} = 46.57 \text{ MeV}$
- Coulomb energy $R_c = 1.24 A^{1/3} \text{ fm}$
- Pairing term must also be considered

Empirical Mass Formula



Plotted: most stable nucleus for a given A

Central density of nuclei

Multiply charge density at the origin by A/Z

\Rightarrow Empirical density = 0.16 nucleons / fm³

\Rightarrow Equivalent to $k_F = 1.33 \text{ fm}^{-1}$

Nuclear Matter

$$N = Z$$

No Coulomb

$A \Rightarrow \infty, V \Rightarrow \infty$ but $A/V = \rho$ fixed

“Two most important numbers”

$$b_{vol} = 15.56 \text{ MeV and } k_F = 1.33 \text{ fm}^{-1}$$

Historical Perspective

- First attempt using scattering in the medium *Brueckner 1954*
- Formal development (linked cluster expansion) *Goldstone 1956*
- Low-density expansion *Galitskii 1958*
- Reorganized perturbation expansion (60s)
involving ordering in the number of hole lines *Bethe & students*
BBG-expansion
- Variational Theory vs. Lowest Order *BBG* (70s) *Clark, Pandharipande*
- Variational results & next hole-line terms (80s) *Day, Wiringa*
- Three-body forces? Relativity? (80s) *Urbana, CUNY*
- Confirmation of three hole-line results (90s) *Baldo et al.*
- New insights from experiment *NIKHEF Amsterdam*
about what nucleons are up to in the nucleus (90s & 00s) *JLab*

Old pain and suffering!

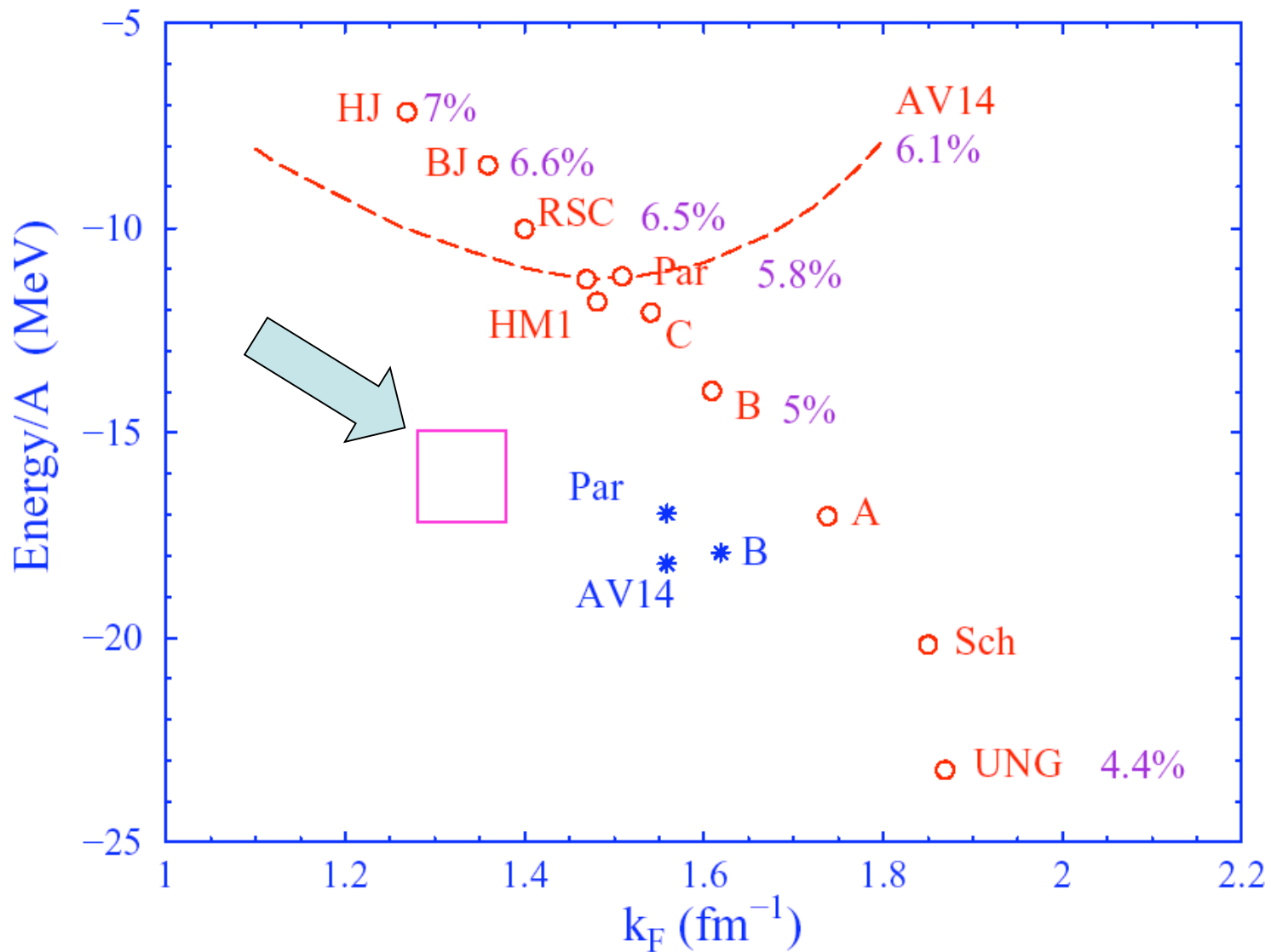
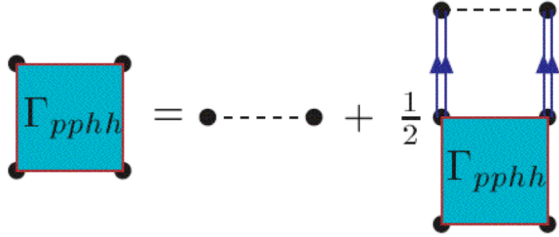


Figure adapted from Marcello Baldo (Catania)

Lowest-order Brueckner theory (two hole lines)



G_{BG}^f angle-average of

$$G_{BG}^f(k_1, k_2; E) = \frac{\theta(k_1 - k_F)\theta(k_2 - k_F)}{E - \varepsilon(k_1) - \varepsilon(k_2) + i\eta}$$

$$\langle k\ell | G^{JST}(K, E) | k'\ell' \rangle = \langle k\ell | V^{JST} | k'\ell' \rangle + \frac{1}{2} \sum_{\ell''} \int_0^\infty \frac{dq}{(2\pi)^3} q^2 \langle k\ell | V^{JST} | q\ell'' \rangle G_{BG}^f(q; K, E) \langle q\ell'' | G^{JST}(K, E) | k'\ell' \rangle$$

Spectrum $\varepsilon_{BHF}(k) = \frac{\hbar^2 k^2}{2m} + \Sigma_{BHF}(k; \varepsilon_{BHF}(k))$

$k < k_F \Rightarrow$ standard choice
all $k \Rightarrow$ continuous choice

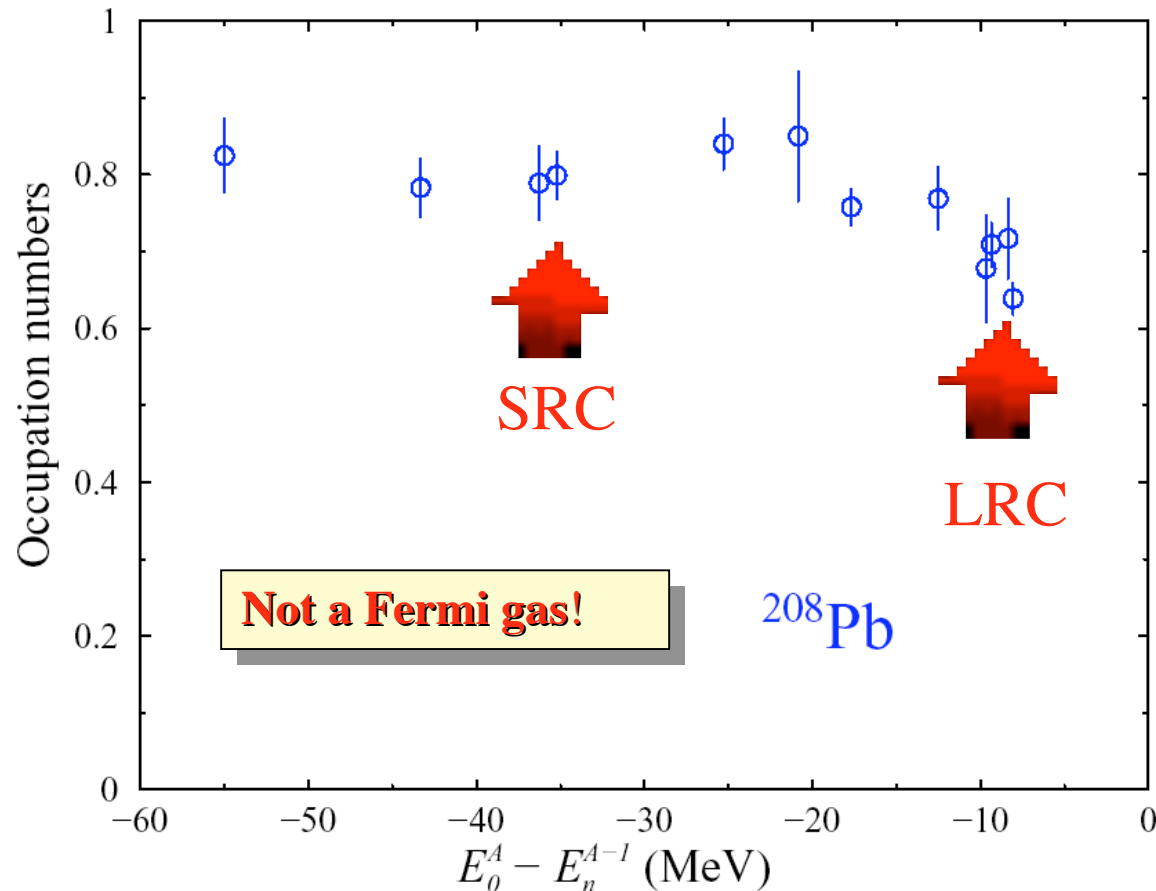
Self-energy $\Sigma_{BHF}(k; E) = \frac{1}{v} \sum_{m, m'} \int \frac{d^3 k'}{(2\pi)^3} \theta(k_F - k') \langle \vec{k}\vec{k}' mm' | G(\vec{k} + \vec{k}'; E + \varepsilon_{BHF}(k')) | \vec{k}\vec{k}' mm' \rangle$

Energy

$$\frac{E}{A} = \frac{4}{\rho} \int \frac{d^3 k}{(2\pi)^3} \theta(k_F - k) \frac{\hbar^2 k^2}{2m} + \frac{1}{2\rho} \sum_{m, m'} \int \frac{d^3 k}{(2\pi)^3} \theta(k_F - k) \int \frac{d^3 k'}{(2\pi)^3} \theta(k_F - k') \langle \vec{k}\vec{k}' mm' | G(\vec{k} + \vec{k}'; \varepsilon_{BHF}(k) + \varepsilon_{BHF}(k')) | \vec{k}\vec{k}' mm' \rangle$$

M. van Batenburg (thesis, 2001) & L. Lapikás from $^{208}\text{Pb} (e, e'p) ^{207}\text{Tl}$

Occupation of deeply-bound proton levels from **EXPERIMENT**



Up to 100 MeV
missing energy
and
270 MeV/c
missing momentum

Covers the whole
mean-field domain
for the FIRST time!!

Confirmation of theory

Where are the last protons?

Answer is coming!

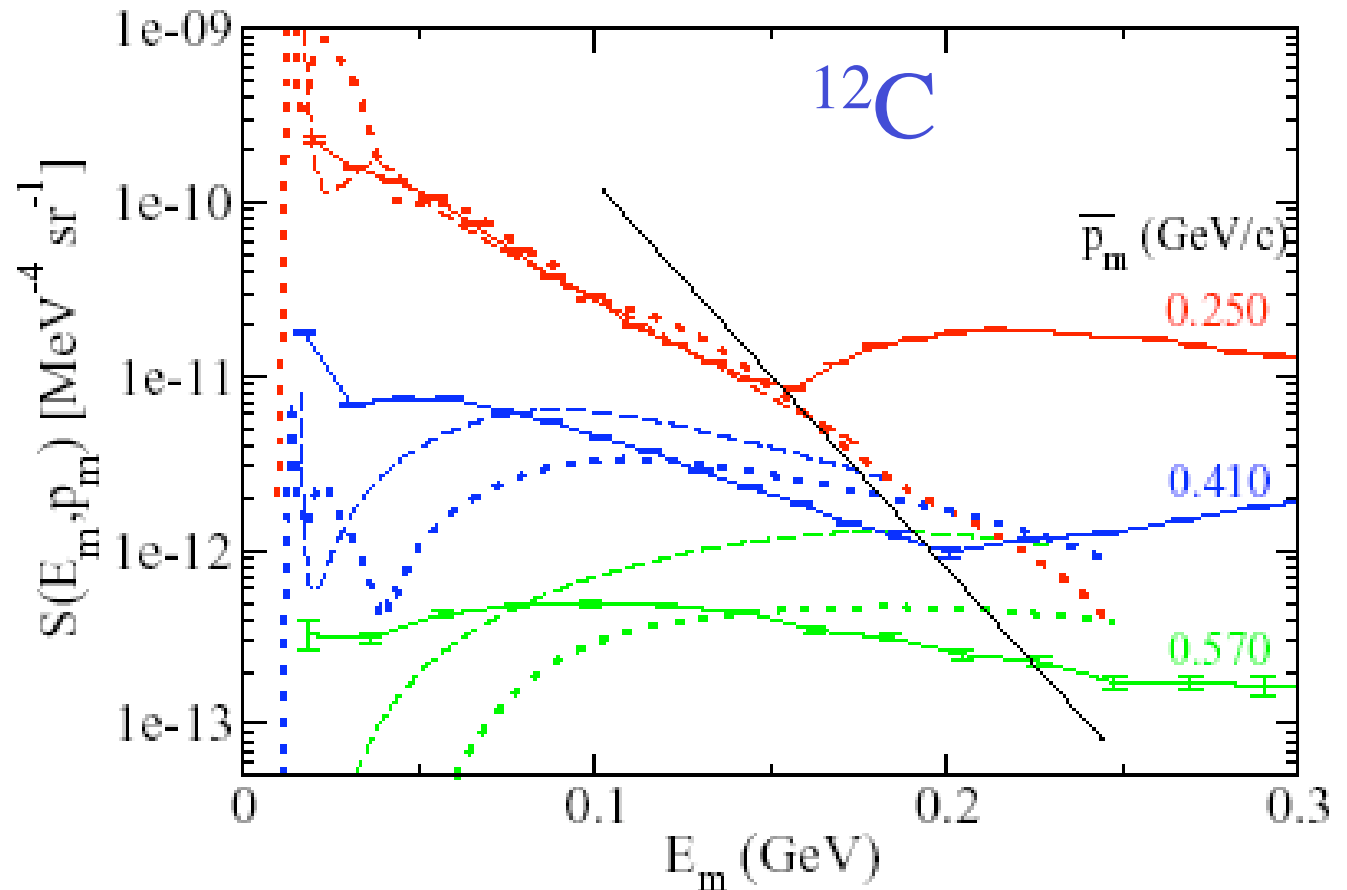
Jlab data

PRL **93**,182501 (2004)

Rohe et al.

Location of high-momentum componen

integrated strength OK!



There are high-momentum components in the nuclear ground state!

Energy Sum Rule (Migdal, Galitskii, Koltun ...)

Finite nuclei

$$E_0^A = \langle \Psi_0^A | \hat{H} | \Psi_0^A \rangle = \frac{1}{2} \sum_{\alpha\beta} \langle \alpha | T | \beta \rangle n_{\alpha\beta} + \frac{1}{2} \sum_{\alpha} \int_{-\infty}^{\varepsilon_F^-} dE E S_h(\alpha; E)$$

$$n_{\alpha\beta} = \langle \Psi_0^A | a_{\alpha}^+ a_{\beta} | \Psi_0^A \rangle = \frac{1}{\pi} \int_{-\infty}^{\varepsilon_F^-} dE \text{Im} G(\beta, \alpha; E)$$

$$S_h(\alpha; E) = \sum_n \left| \langle \Psi_n^{A-1} | a_{\alpha} | \Psi_0^A \rangle \right|^2 \delta(E - (E_0^A - E_n^{A-1})) = \frac{1}{\pi} \text{Im} G(\alpha, \alpha; E)$$

Nuclear matter

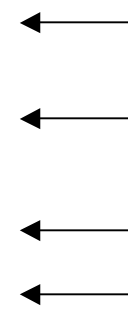
$$\frac{E}{A} = \frac{1}{2} \left\{ \frac{4}{\rho} \int \frac{d^3 k}{(2\pi)^3} \int_{-\infty}^{\varepsilon_F} dE \left(\frac{\hbar^2 k^2}{2m} + E \right) S_h(k; E) \right\}$$

- Presumes only two-body interactions!
- Correct description of experimental spectral function should yield good E/A !!

Where does binding come from (really)?

¹⁶O PRC51,3040(1995)

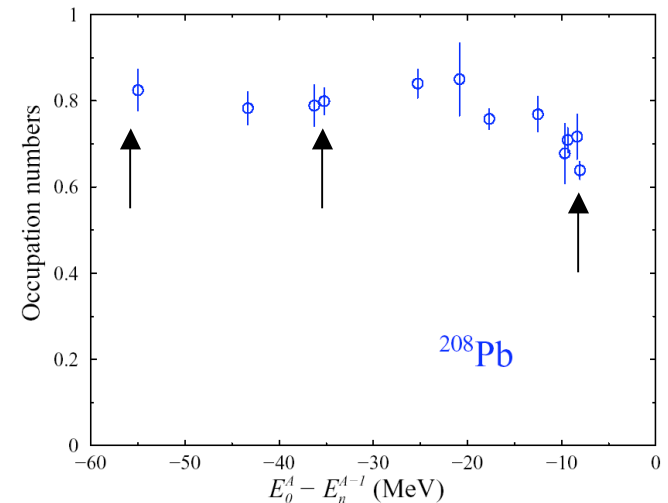
lj	"BHF"			Total		
	ϵ	t	ΔE	ϵ	t	ΔE
$s_{\frac{1}{2}}$ qh	-36.9	11.8	-50.3	-34.3	11.2	-36.0
$s_{\frac{1}{2}}$ c				-90.4	17.1	-22.9
$p_{\frac{3}{2}}$ qh	-15.4	17.6	9.1	-17.9	18.1	0.4
$p_{\frac{3}{2}}$ c				-95.2	35.2	-10.0
$p_{\frac{1}{2}}$ qh	-11.5	16.6	10.3	-14.1	17.2	5.5
$p_{\frac{1}{2}}$ c				-103.6	35.9	-5.8
$\ell > 1$ c				-98.9	63.2	-12.3
$E/A(\text{MeV})$		-1.9			-5.1	
$\langle r \rangle(\text{fm})$		2.59			2.55	



Quasiholes contribute 37% to the total energy
 High-momentum nucleons (continuum) contribute 63%
 but represent **only** about 10% of the particles!!

Saturation density and SRC

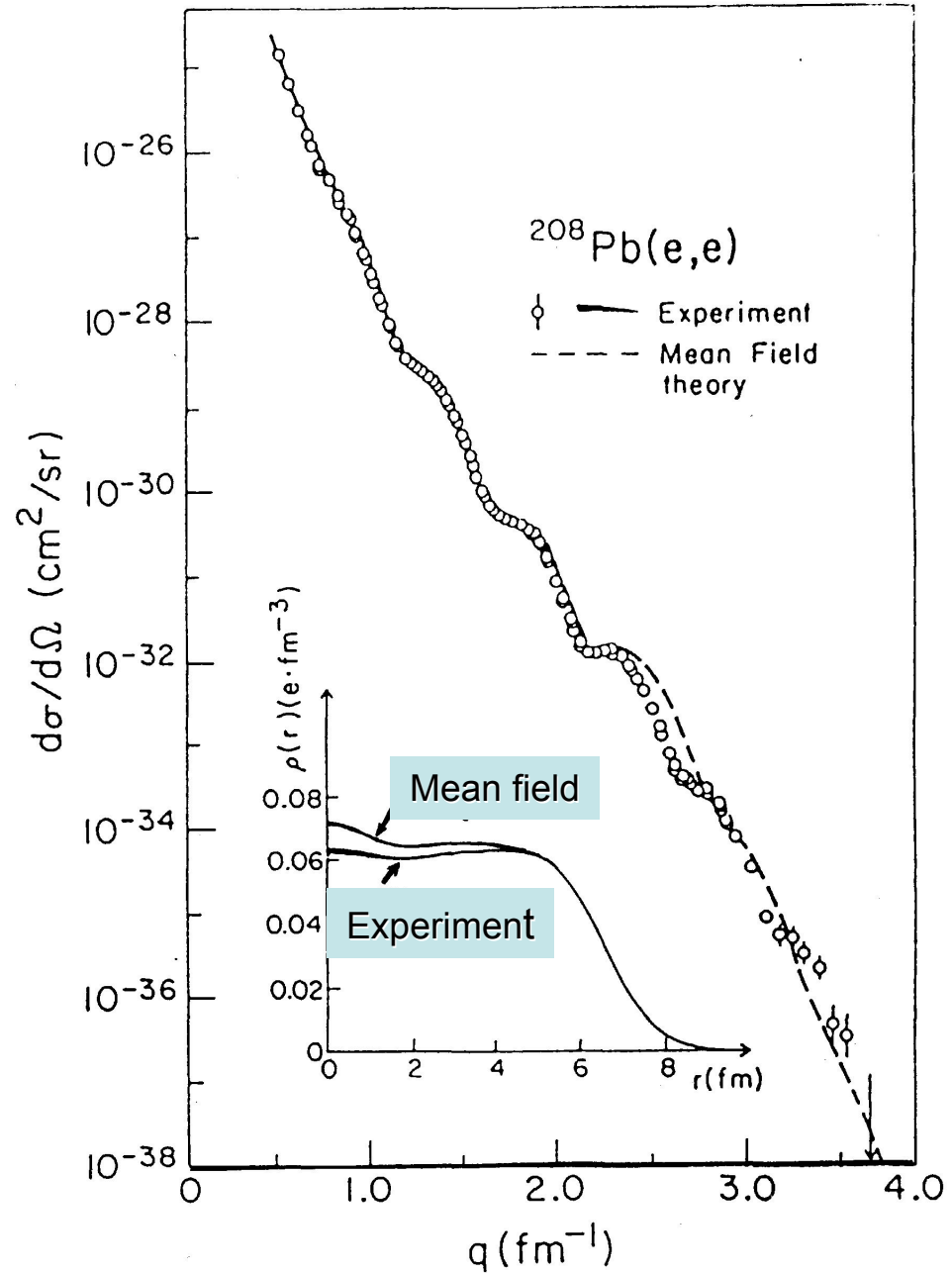
- Saturation density related to nuclear charge density at the origin. Data for $^{208}\text{Pb} \Rightarrow A/Z * \rho_{\text{ch}}(0) = 0.16 \text{ fm}^{-3}$
- Charge at the origin determined by protons in s states
- Occupation of $0s$ and $1s$ totally dominated by SRC as can be concluded from recent analysis of $^{208}\text{Pb}(e,e'p)$ data and theoretical calculations of occupation numbers in nuclei and nuclear matter.
- Depletion of $2s$ proton also dominated by SRC:
15% of the total depletion of 25% ($n_{2s} = 0.75$)



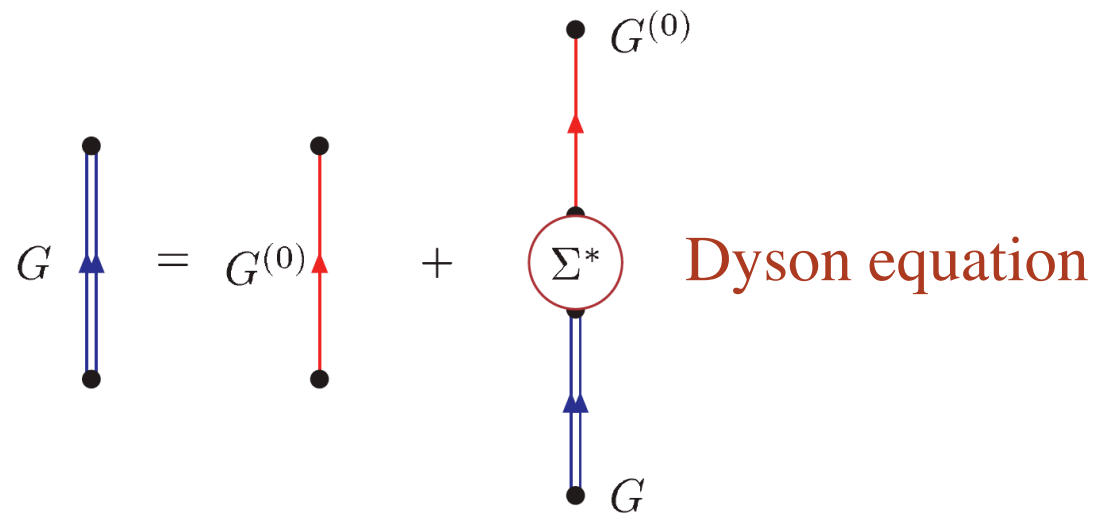
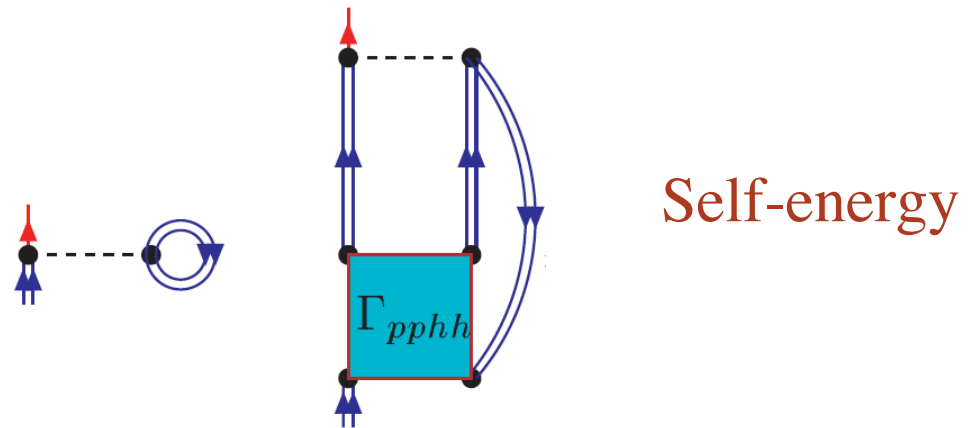
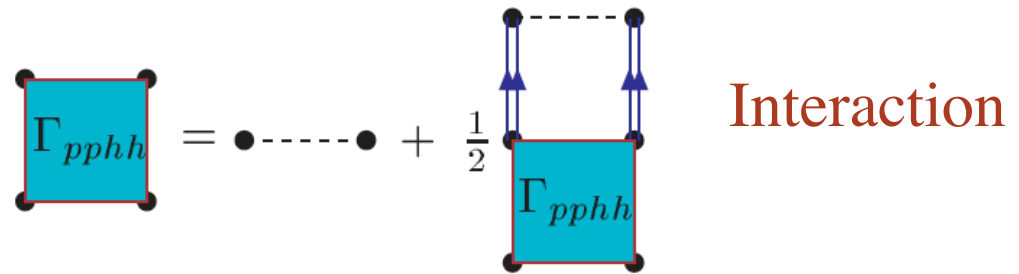
- Conclusion: **Nuclear saturation dominated by SRC
and therefore high-momentum components**

Elastic electron scattering from ^{208}Pb

B. Frois *et al.*
Phys. Rev. Lett. **38**, 152 (1977)



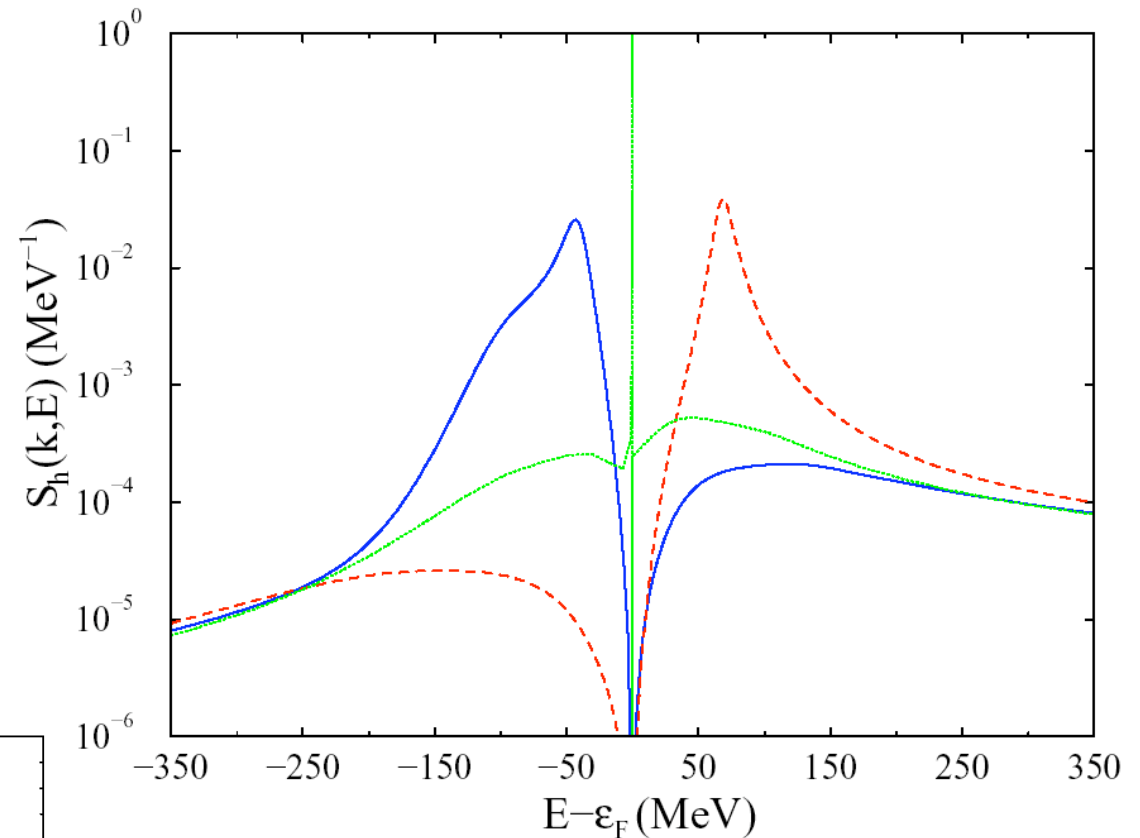
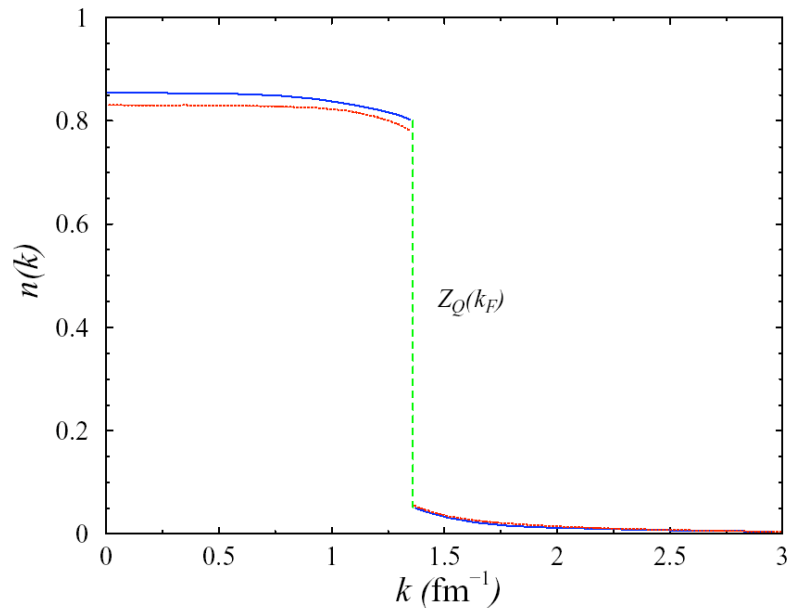
Self-consistent treatment of SRC in nuclear matter



Results from Nuclear Matter

2nd generation (2000)

- Spectral functions for $k = 0, 1.36, \& 2.1 \text{ fm}^{-1}$
- Common tails on both sides of ϵ_F



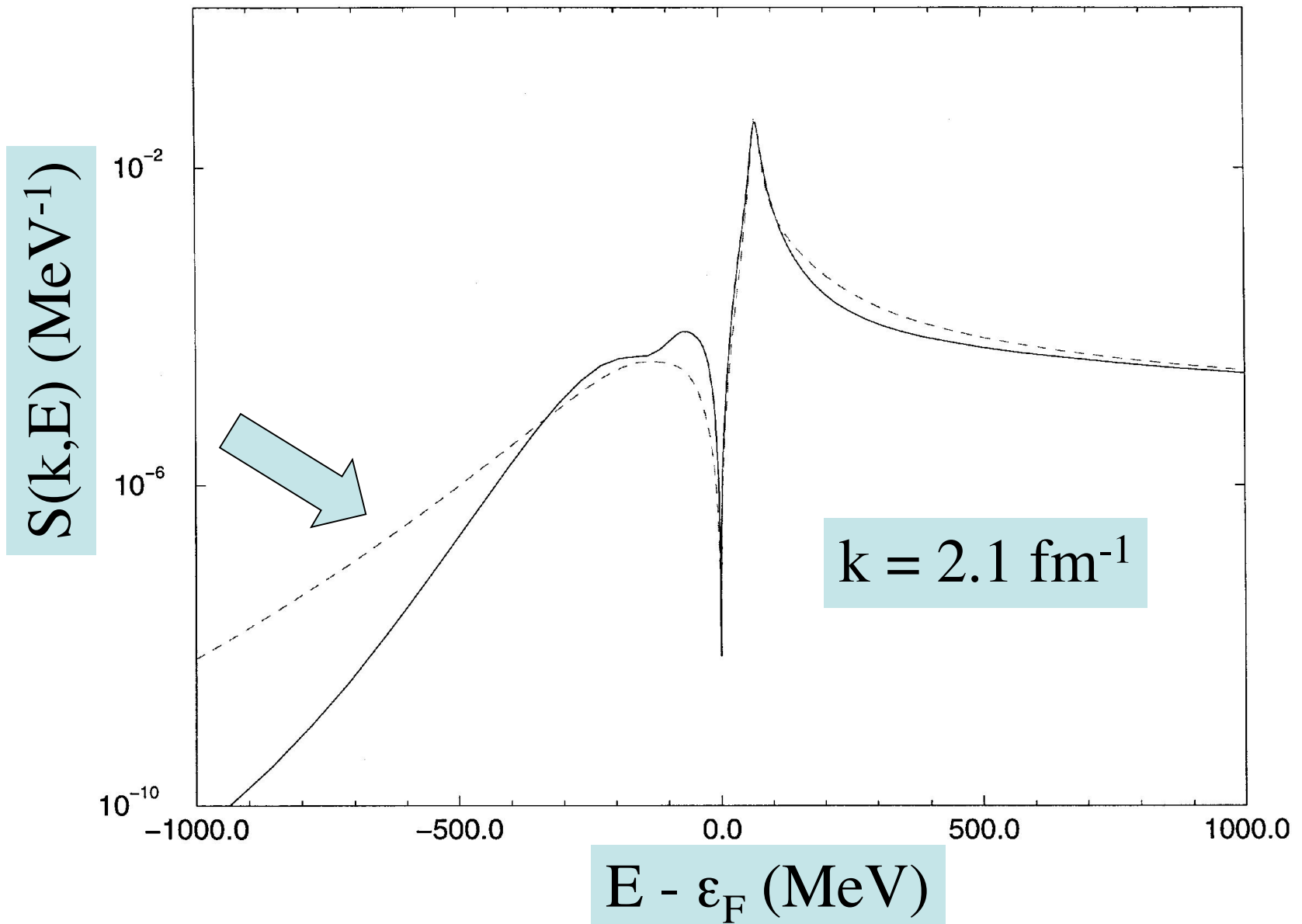
Momentum distribution :

only minor changes

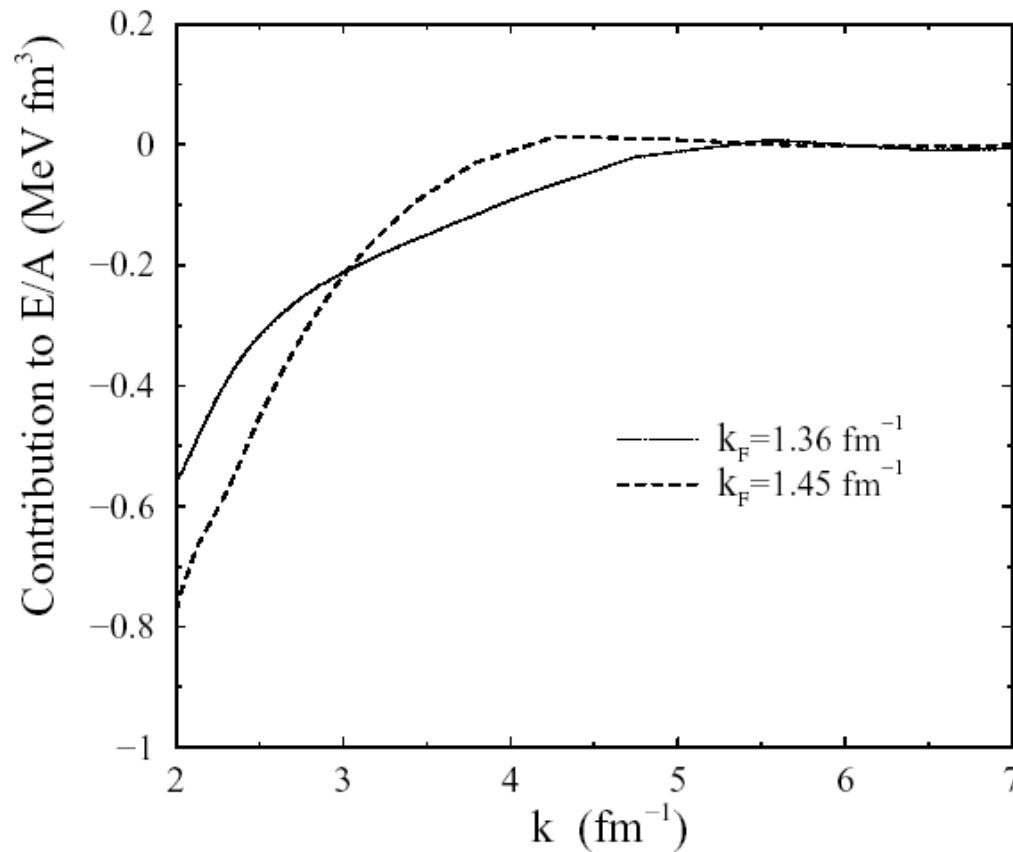
occupation in nuclei

depleted similarly!?!

Self-consistent spectral functions



Saturation with self-consistent spectral functions in nuclear matter \Rightarrow reasonable saturation properties

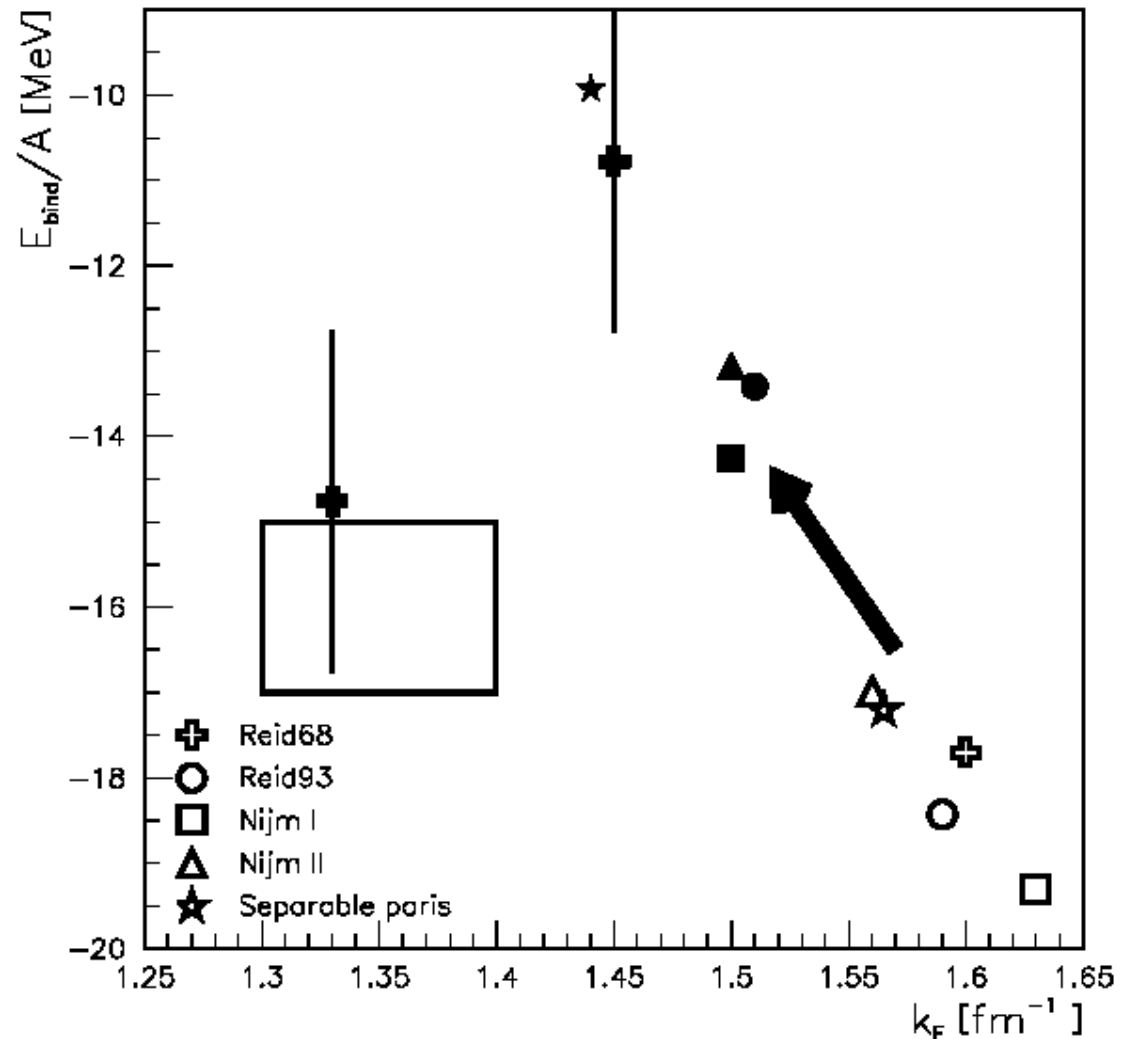


Contribution to the energy per particle before integration over the single-particle momentum at high momentum for two densities

Saturation of Nuclear Matter

Ladders and self-consistency for Nuclear Matter

- Ghent group
Dewulf, Van Neck &
Waroquier
- St. Louis
Stoddard, WD



Phys. Rev. Lett. **90**, 152501 (2003)

Self-consistent spectral functions

- Distribution below ε_F broadens for high momenta and develops a common tail at high missing energy
- Slight increase in occupation $k < k_F$ to 85% at $k_F = 1.36 \text{ fm}^{-1}$ compared to Phys. Rev. C44, R1265 (1991) & Nucl. Phys. A555, 1 (1993)
- Self-consistent treatment of Pauli principle
- Interaction between dressed particles weaker (reduced cross sections for both pn and nn)
- Pairing instabilities disappear in all channels
- Saturation with lower density than before and reasonable binding
- **Contribution of long-range correlations excluded**

Self-consistent Green's functions and the energy of the ground state of the electron gas



GW approximation

G self-consistent sp propagator

W screened Coulomb interaction

⇒ RPA with dressed sp propagators

Electron gas : -XC energies (Hartrees)

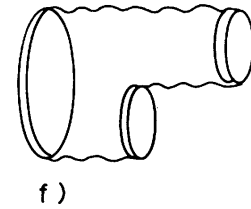
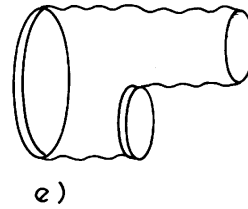
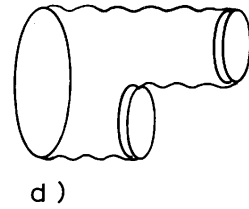
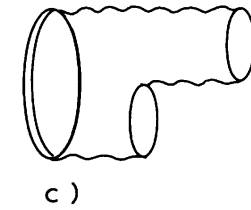
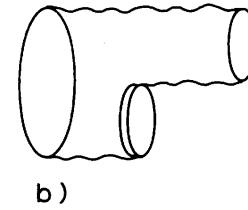
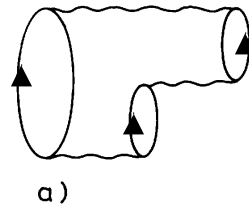
<u>Method</u>	$r_s = 1$	$r_s = 2$	$r_s = 4$	$r_s = 5$	$r_s = 10$	$r_s = 20$	Reference
QMC	0.5180	0.2742	0.1464	0.1197	0.0644	0.0344	CA80
	0.5144	0.2729	0.1474	0.1199	0.0641	0.0344	OB94;OHB99
GW	0.5160	0.2727	0.1450	0.1185	0.0620	0.032	GG01
		0.2741	0.1465				HB98
RPA	0.5370	0.2909	0.1613	0.1340	0.0764	0.0543	

What about long-range correlations in nuclear matter?

- Collective excitations in nuclei very different from those in nuclear matter
- Long-range correlations normally associated with small q
- Contribution to the energy like $dq q^2 \Rightarrow$ very small (except for e-gas)
- Contributions of collective excitations to the binding energy of nuclear matter dominated by pion-exchange induced excitations?!?

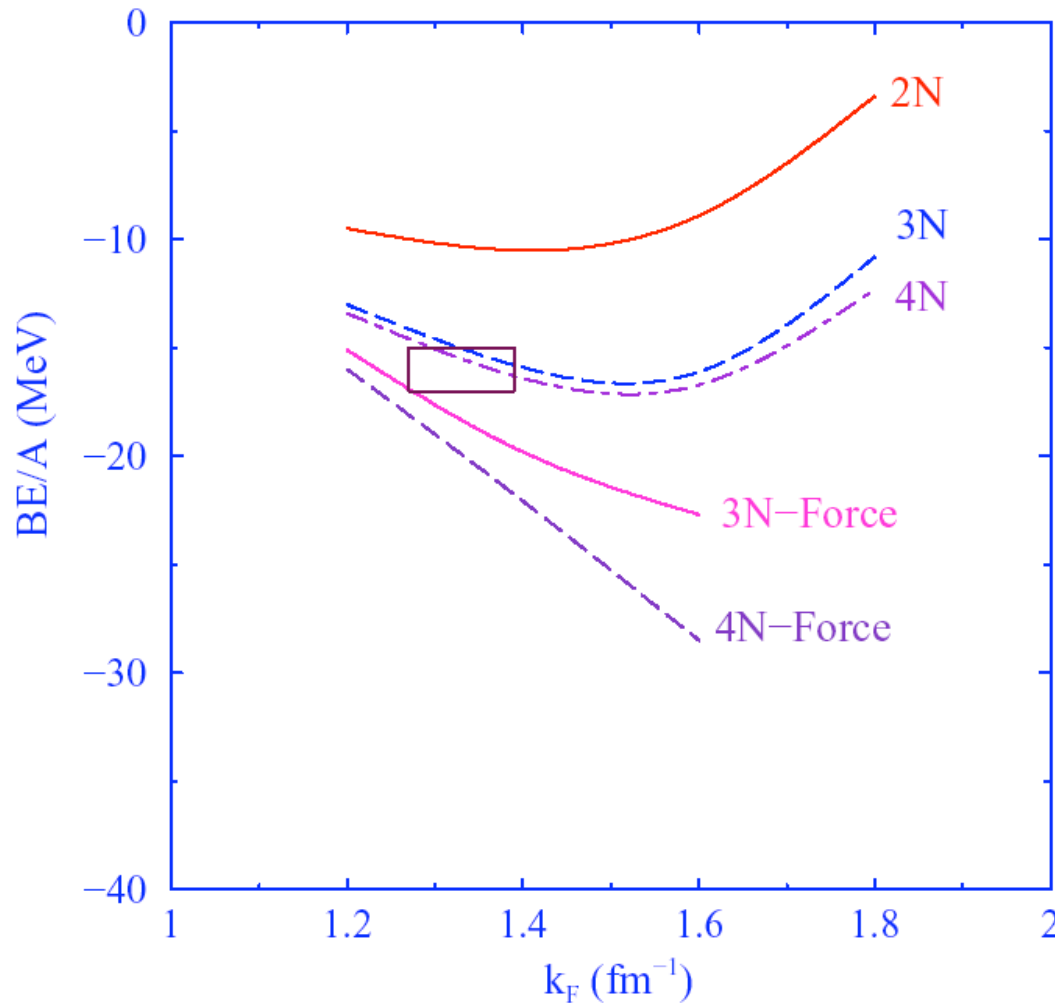
Inclusion of Δ - isobars as “3N-” and “4N-force”

Nucl. Phys. A389, 492 (1982)



k_F [fm ⁻¹]	1.0	1.2	1.4	1.6
third order				
a)	-0.303	-1.269	-3.019	-5.384
b)	-0.654	-1.506	-2.932	-5.038
c)	-0.047	-0.164	-0.484	-1.175
d)	0.033	0.095	0.220	0.447
e)	-0.104	-0.264	-0.589	-1.187
f)	0.041	0.137	0.385	0.962
Sum	-1.034	-2.971	-6.419	-11.375

Inclusion of Δ -isobars as 3N- and 4N-force



2N,3N, and 4N from
B.D.Day, PRC24,1203(81)

Rings with Δ -isobars :

Nucl. Phys. A389, 492 (1982)

PPNPhys 12, 529 (1983)

\Rightarrow No sensible convergence with Δ -isobars

Nuclear Saturation without π -collectivity

- Variational calculations treat LRC (on average) and SRC simultaneously (Parquet equivalence) so **difficult** to separate LRC and SRC
- Remove 3-body ring diagram from Catania hole-line expansion calculation \Rightarrow about the correct saturation density
- Hole-line expansion doesn't treat Pauli principle very well
- Present results improve treatment of Pauli principle by self-consistency of spectral functions \Rightarrow more reasonable saturation density and binding energy acceptable
- Neutron matter: pionic contributions must be included (Δ)

Pion collectivity: nuclei vs. nuclear matter

- Pion collectivity leads to pion condensation at higher density in nuclear matter (including Δ -isobars) \Rightarrow Migdal ...
- No such collectivity observed in nuclei \Rightarrow LAMPF / Osaka data
- Momentum conservation in nuclear matter dramatically favors amplification of π -exchange interaction at fixed q
- In nuclei the same interaction is sampled over all momenta Phys. Lett. **B146**, 1(1984)

$$V_{\pi}(q) = -\frac{f_{\pi}^2}{m_{\pi}^2} \frac{q^2}{m_{\pi}^2 + q^2}$$

Needs further study

\Rightarrow Exclude collective pionic contributions to nuclear matter binding energy

Two Nuclear Matter Problems

The usual one

- With π -collectivity and only nucleons
- Variational + CBF and three hole-line results presumed OK (for E/A) but not directly relevant for comparison with nuclei!
- **NOT OK** if Δ -isobars are included
- Relevant for neutron matter

The relevant one?!

- Without π -collectivity
- Treat only SRC
- But at a sophisticated level by using self-consistency
- Confirmation from Ghent results \Rightarrow Phys. Rev. Lett. **90**, 152501 (2003)
- 3N-forces difficult $\Rightarrow \pi$...
- Relativity?

Comments

Relativity

- Saturation depends on NN σ -coupling in medium but underlying correlated two-pion exchange behaves differently in medium
- $m^* \rightarrow 0$ with increasing ρ opposite in liquid ${}^3\text{He}$ appears unphysical
- Dirac sea under control?
- sp strength overestimated too many nucleons for $k < k_F$

Three-body forces

- Microscopic models yield only attraction in matter and more so with increasing ρ
- Microscopic background of phenomenological repulsion in 3N-force (if it exists)?
- 4N-, etc. forces yield increasing attraction with ρ
- Needed in light nuclei and attractive!
- Mediated by π -exchange
- Argonne group can't get nuclear matter right with new 3N-force

Conclusions

- Good understanding of role of short-range correlations
 - Depletion of Fermi sea: nuclear matter OK for nuclei
 - Confirmed by experiment
 - High-momentum components
 - # of protons experimentally confirmed
- Long-range correlations crucial for distribution of sp strength
- Energy per particle from self-consistent Green's functions
- Better understanding of nuclear matter saturation
⇒ SRC dominate (don't treat LRC from pions)
- We know what protons are up to in nuclei!!